

# A new approach for an integrated kinematic-dynamic orbit determination of low flying satellites based on GNSS observations

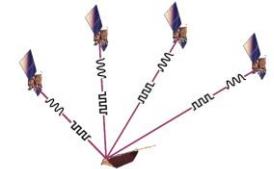
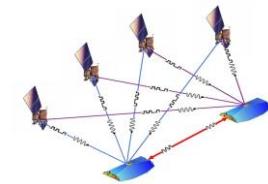
Akbar Shabanlou

Geodätische Woche 2007

Session GNSS (G6)

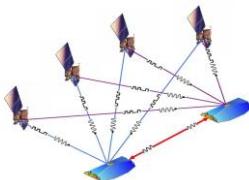
Leipzig, Deutschland

26 Sep. 2007

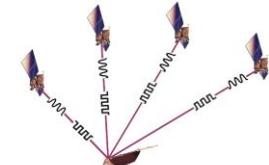


# LEO Missions (Earth Explorers)





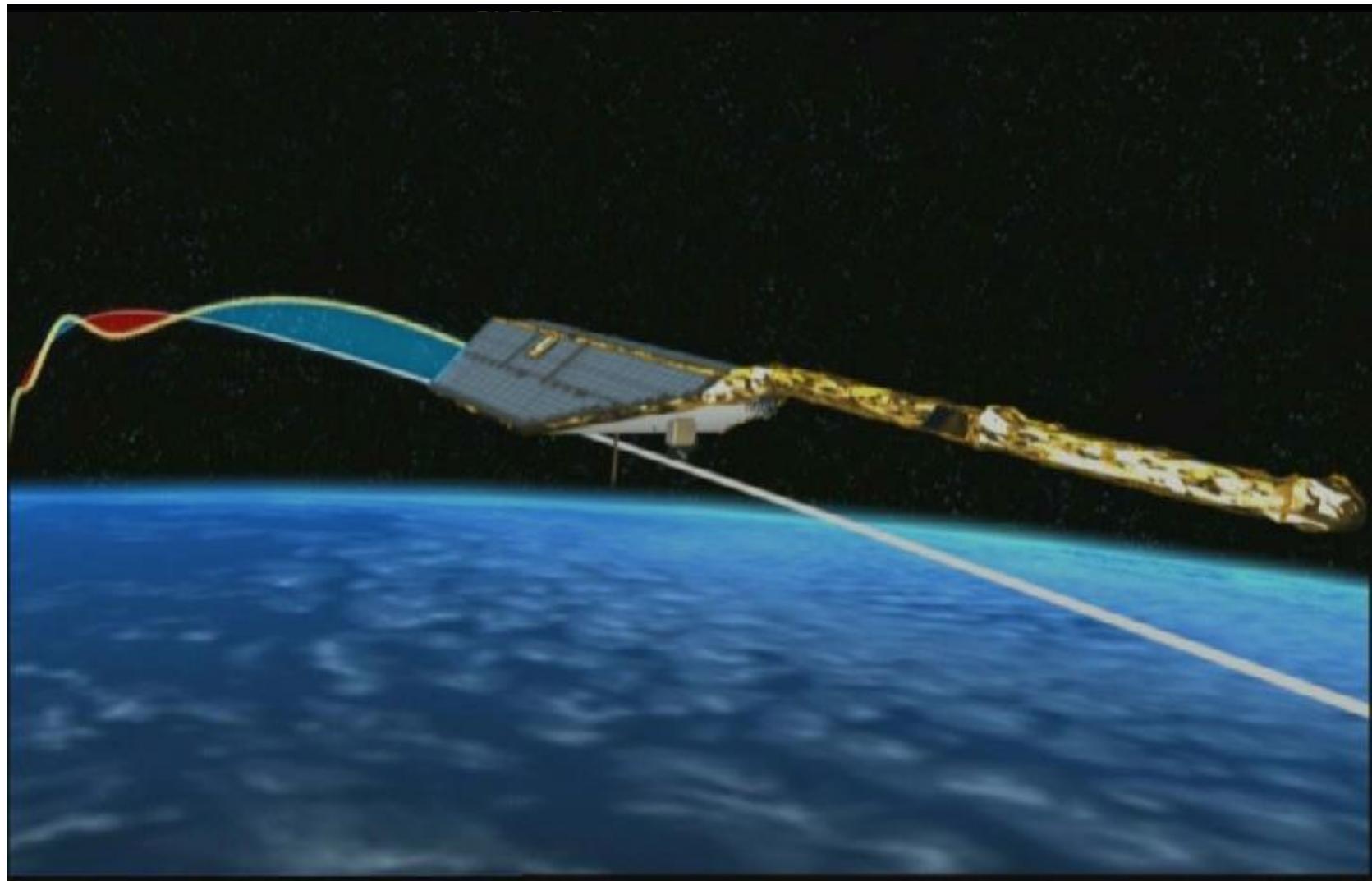
## Advantages of Kinematical LEO Precise Orbit Determination (POD)

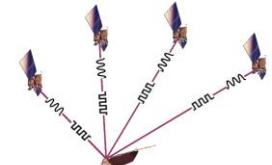
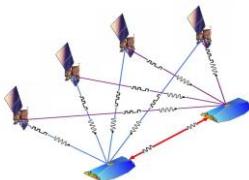


- ✓ Kinematical LEO orbits can be used to recover the gravity field of the Earth by the POD methods,
- ✓ Analysis of altimetry observations requires kinematical orbits of altimetry satellite,
- ✓ Atmosphere sounding requires kinematical positions of the LEO satellites
- ✓ GNSS (GPS, GLONASS, GALILEO) methods play an important role in geometrical (subsequently kinematical) POD in addition to classical methods (e.g. SLR...)



# Precise Orbit Determination (POD)





## Principal techniques of POD

### ✓ Geometrical orbit determination

Only geometrical observations are used, no force models and no constraints; pointwise representation



### ✓ Kinematical orbit determination

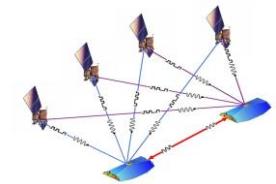
Geometrical & kinematical observations are used, no force models used; representation by kinematical functions



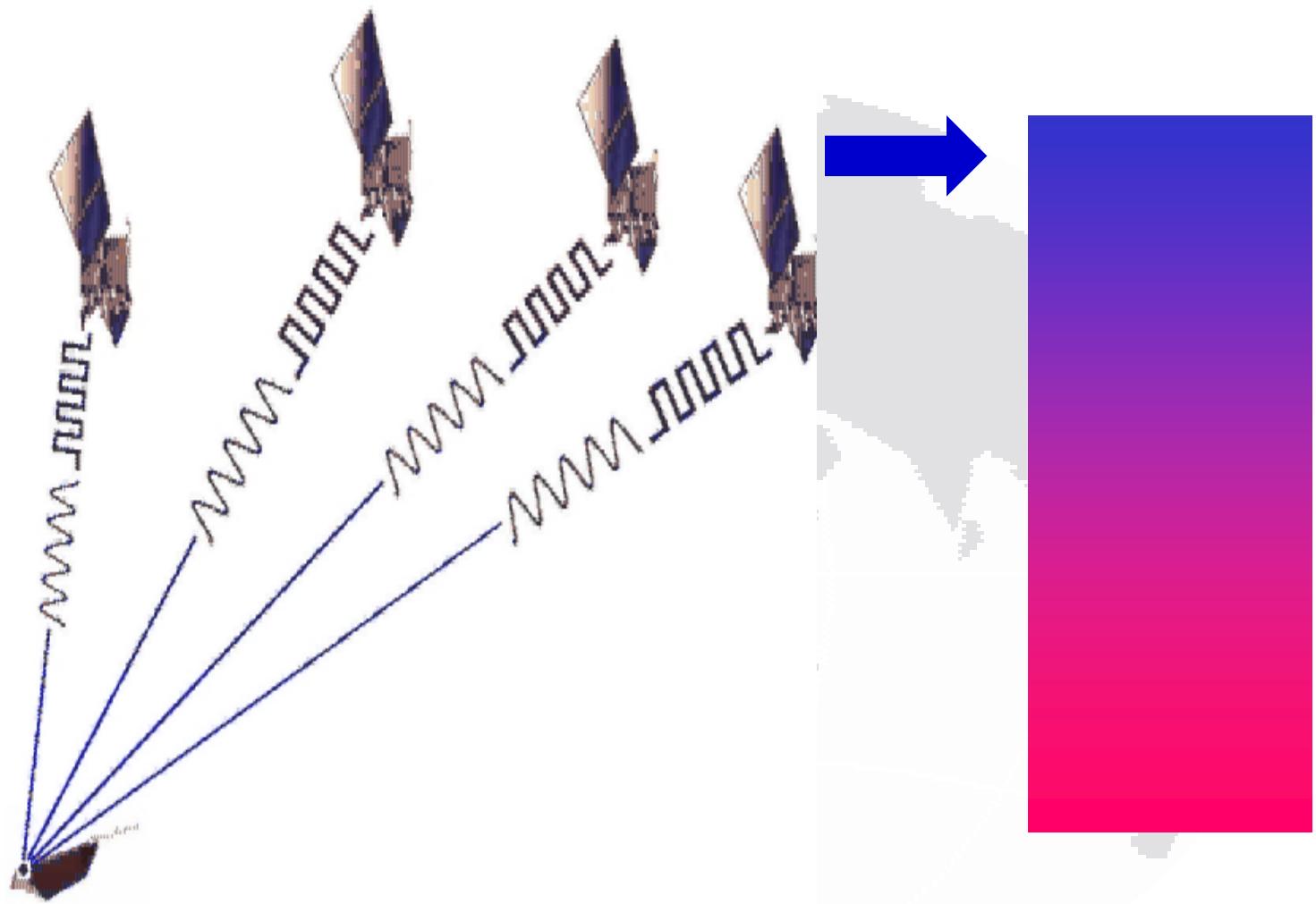
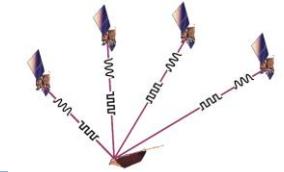
### ✓ Dynamical orbit determination

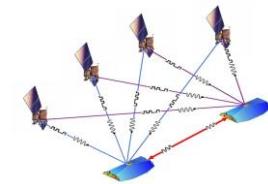
Geometrical, kinematical & dynamical observations are used, force models necessary; pointwise representation or representation by functions



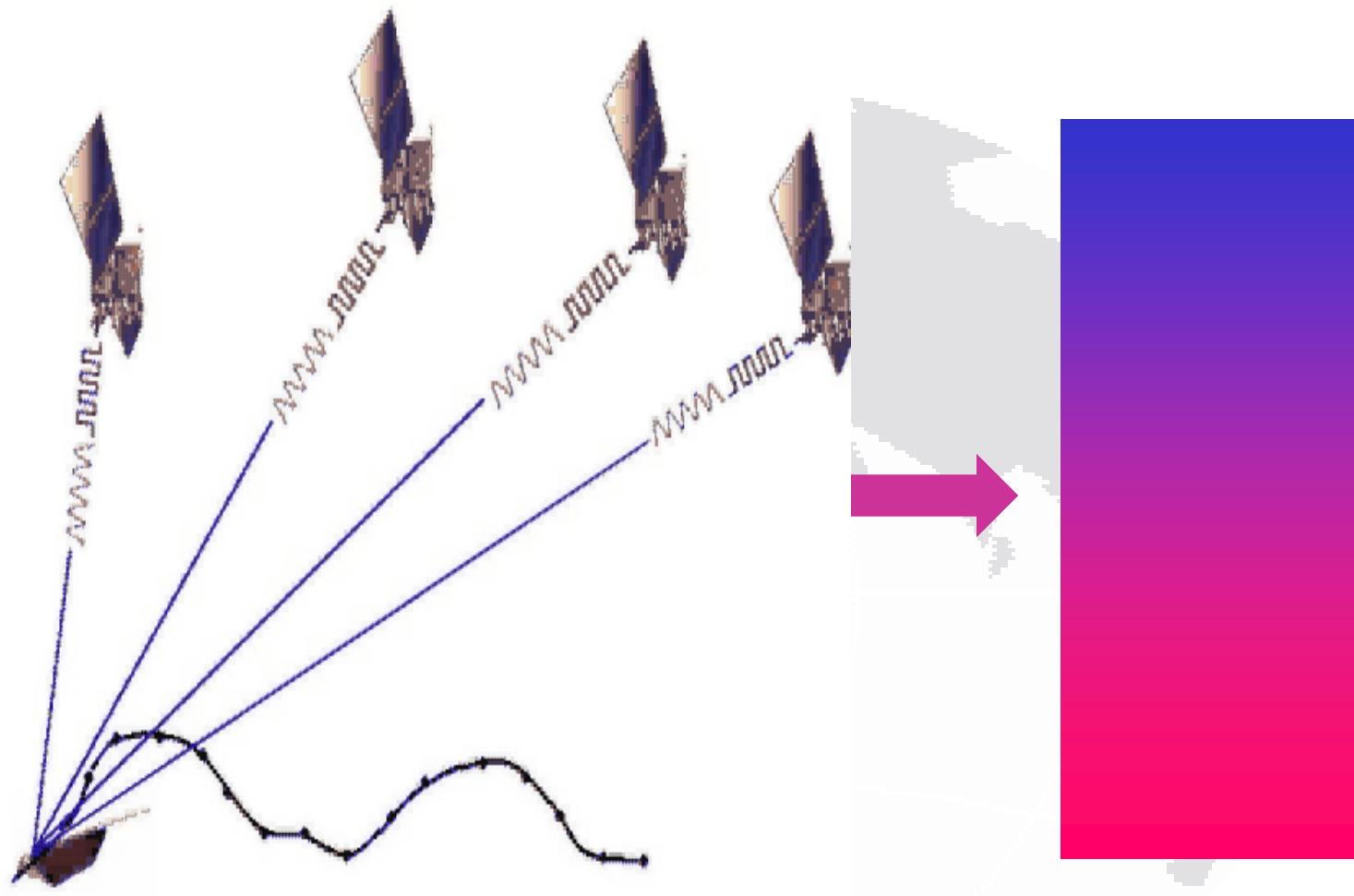
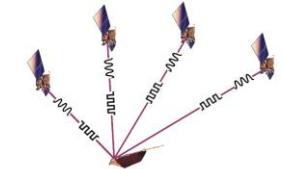


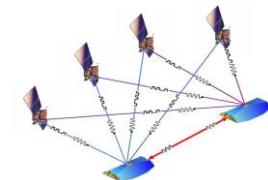
# Geometrical precise orbit determination



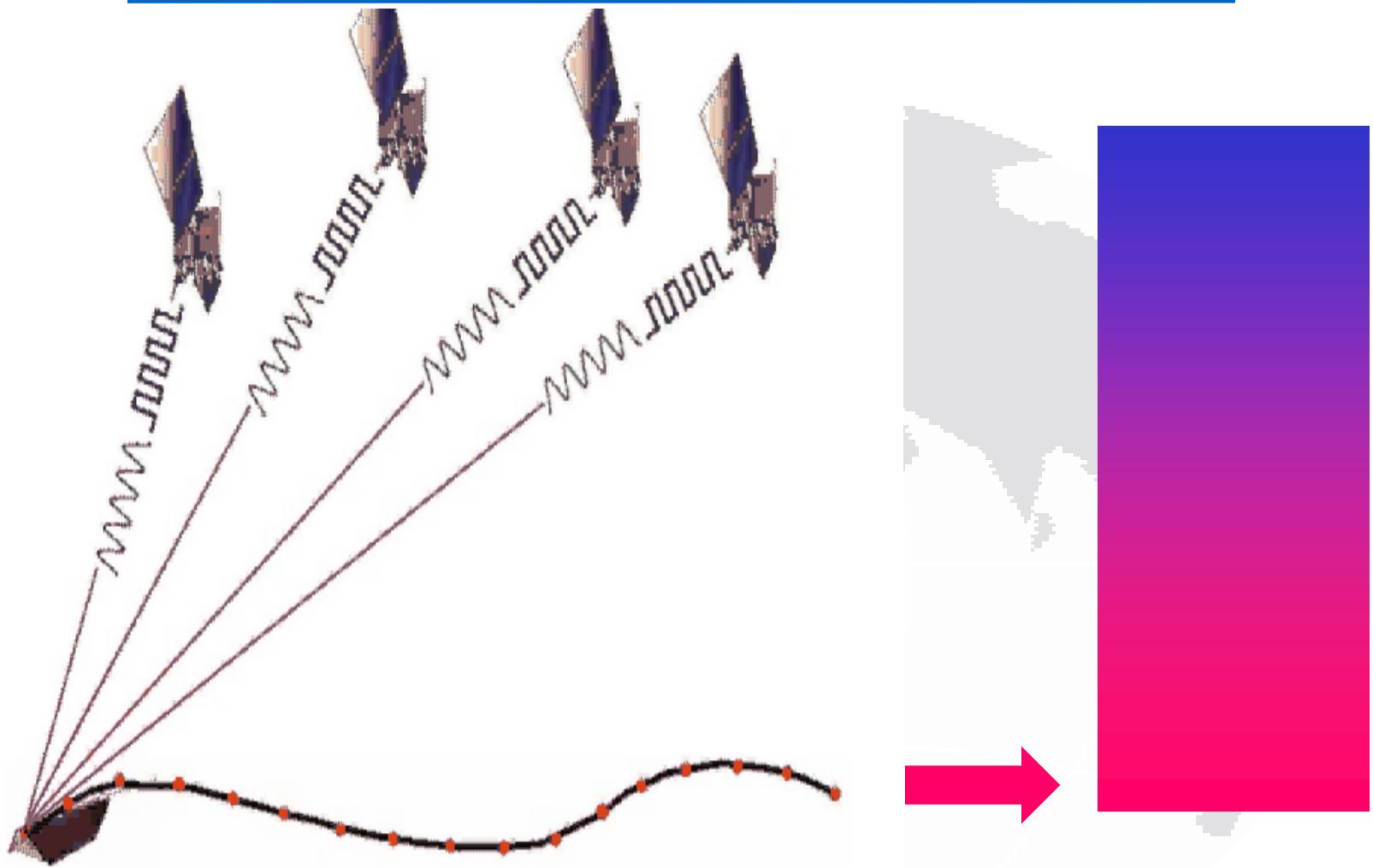
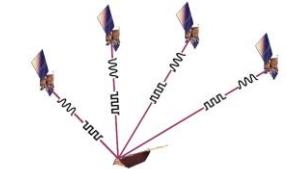


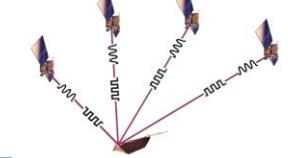
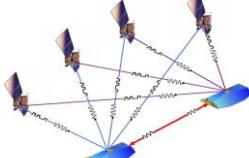
# Kinematical precise orbit determination





# Dynamical precise orbit determination





# Processing concepts

Code measurements

Phase measurements

Zero differencing concept

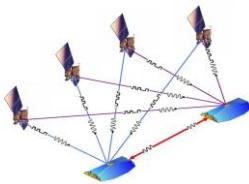
Geometrical orbit  
determination

Kinematical orbit  
determination

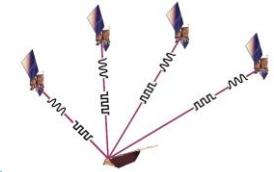
Dynamical orbit  
determination



# Kinematical (short arcs) POD concept

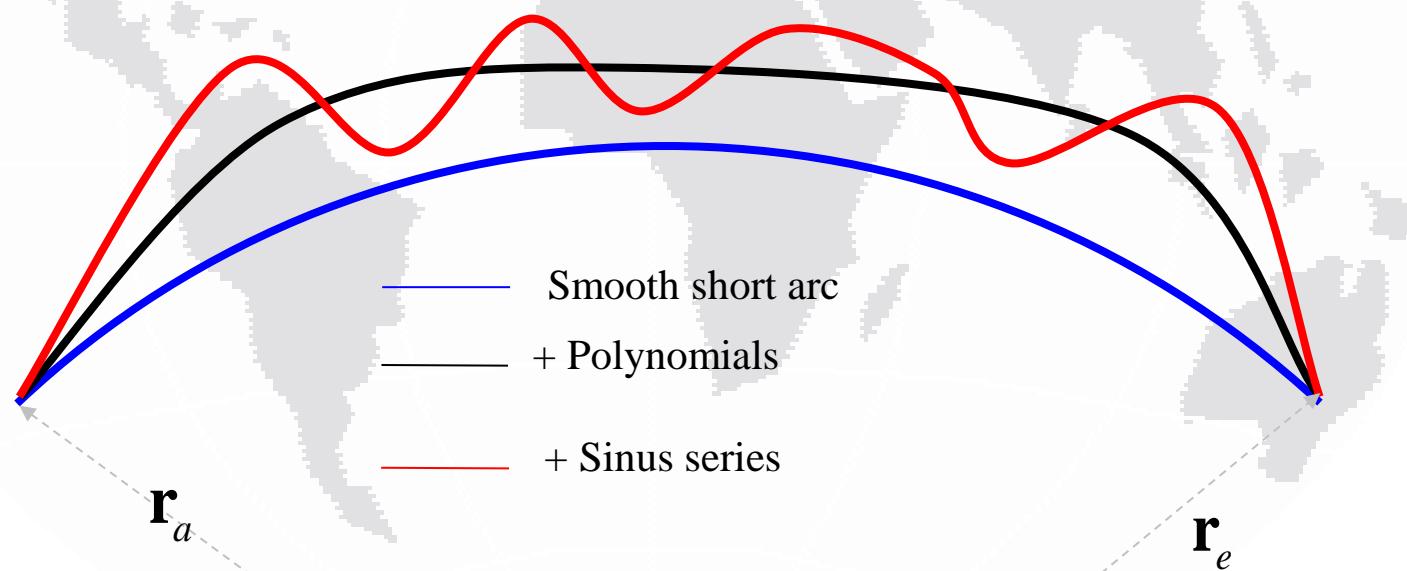


## LEO short arc POD principle



Step by step presentation of LEO short arcs from the solution of Newton-Euler motion equation,

$$\mathbf{r} = \underbrace{\mathbf{r}_a \cdot \frac{\sin((1-\tau) \cdot N)}{\sin(N)} + \mathbf{r}_e \cdot \frac{\sin(\tau N)}{\sin(N)}}_{\text{Smooth short arc}} + \underbrace{\mathbf{C}^T \mathbf{P}(\tau)}_{\text{Euler-Bernoulli Polynomial}} + \underbrace{\sum_{v=1}^n \bar{\mathbf{d}}_v \sin(v\pi\tau)}_{\text{Sinus series}}$$



# Kinematical short arc POD-simulation



selected a-priori

$$\mathbf{r}(t) = \mathbf{r}_a \cdot \frac{\sin((1-\tau)N)}{\sin(N)} + \mathbf{r}_e \cdot \frac{\sin(\tau N)}{\sin(N)} + \mathbf{C}^T \mathbf{P}(\tau) + \sum_{v=1}^n \bar{\mathbf{d}}_v \sin(v\pi\tau)$$



all coefficients can be estimated by a Gauss-Markov model

Advantage:

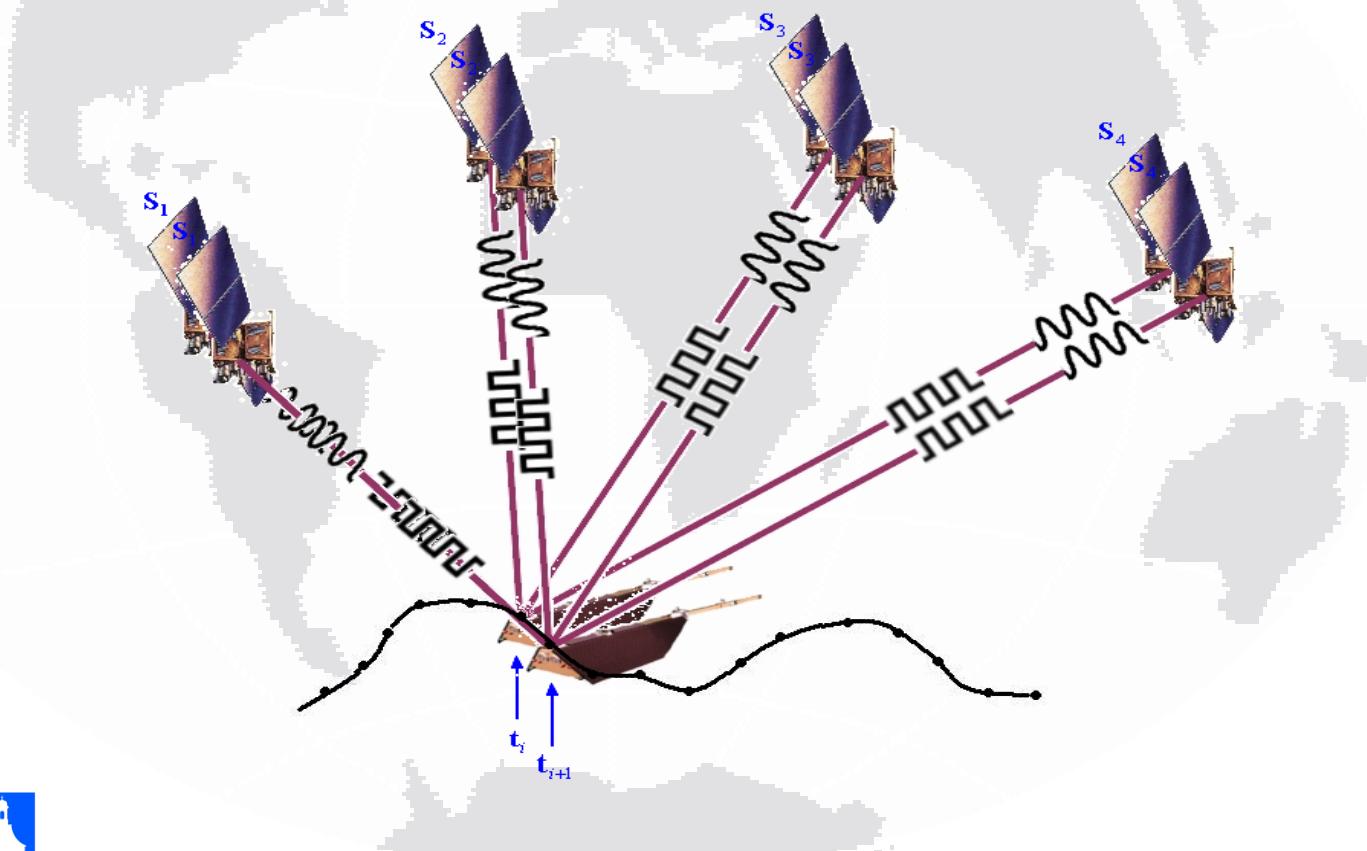
the kinematical orbits and other kinematical parameters can be derived directly from estimated LEO short arc parameters.

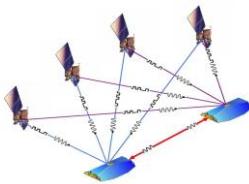


# Kinematical (short arcs) POD simulated case

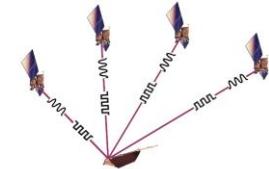
# Absolute position from sequential time differenced carrier phase

- ✓ cut-off angle 15°
- ✓ simple data screening & S/N filtering
- ✓ elevation weighting



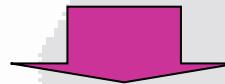


## Kinematical short arc POD sequential time differenced carrier phase



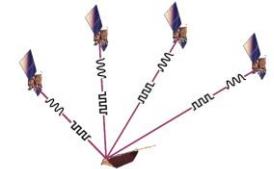
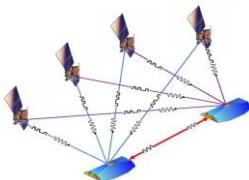
Sequential time differenced carrier phase observations between GPS satellite  $s$  LEO receiver  $r$  at epoch  $t$  can be written as:

$$\Delta\Phi_r^s(t_1, t_2) = |\mathbf{r}^s(t_2) - \mathbf{r}_r(t_2)| - |\mathbf{r}^s(t_1) - \mathbf{r}_r(t_1)| + e_{\Delta\Phi_3}$$



$$\Delta\Phi_{r,3}^s(t_1, t_2) =$$

$$\left| \mathbf{r}^s(t_2) - \left( \mathbf{r}_a \frac{\sin((1-\tau_2)N)}{\sin(N)} + \mathbf{r}_e \frac{\sin(\tau_2 N)}{\sin(N)} + \mathbf{C}_{3 \times 4} \mathbf{P}(\tau_2) + \sum_{f=1}^n \bar{\mathbf{d}}_f \sin(\pi f \tau_2) \right) \right| - \left| \mathbf{r}^s(t_1) - \left( \mathbf{r}_a \frac{\sin((1-\tau_1)N)}{\sin(N)} + \mathbf{r}_e \frac{\sin(\tau_1 N)}{\sin(N)} + \mathbf{C}_{3 \times 4} \mathbf{P}(\tau_1) + \sum_{f=1}^n \bar{\mathbf{d}}_f \sin(\pi f \tau_1) \right) \right| + e_{\Delta\Phi_3}$$



## Gauss-Markov model

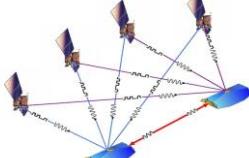
The Gauss-Markov model for two sub-sequential epochs,

$$\underbrace{\begin{pmatrix} \Delta\Delta\Phi_r^{s_1}(t_1, t_2) \\ \vdots \\ \Delta\Delta\Phi_r^{s_j}(t_1, t_2) \\ \vdots \\ \Delta\Delta\Phi_r^{s_{m_1}}(t_1, t_2) \end{pmatrix}}_{\text{Residuals}} = \underbrace{\begin{pmatrix} \mathbf{A}_{\mathbf{r}_r}^{s_1}(t_2)\mathbf{A}_{\mathbf{X}_r}^{\mathbf{r}_r}(t_2) - \mathbf{A}_{\mathbf{r}_r}^{s_1}(t_1)\mathbf{A}_{\mathbf{X}_r}^{\mathbf{r}_r}(t_1) \\ \vdots \\ \mathbf{A}_{\mathbf{r}_r}^{s_j}(t_2)\mathbf{A}_{\mathbf{X}_r}^{\mathbf{r}_r}(t_2) - \mathbf{A}_{\mathbf{r}_r}^{s_j}(t_1)\mathbf{A}_{\mathbf{X}_r}^{\mathbf{r}_r}(t_1) \\ \vdots \\ \mathbf{A}_{\mathbf{r}_r}^{s_{m_1}}(t_2)\mathbf{A}_{\mathbf{X}_r}^{\mathbf{r}_r}(t_2) - \mathbf{A}_{\mathbf{r}_r}^{s_{m_1}}(t_1)\mathbf{A}_{\mathbf{X}_r}^{\mathbf{r}_r}(t_1) \end{pmatrix}}_{\text{Design matrix}} \underbrace{[\mathbf{X}_r - \mathbf{X}_r^0]}_{\text{Unknowns}}$$

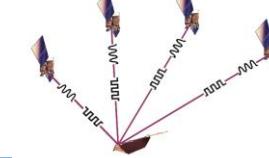
$$\mathbf{X}_r = \begin{pmatrix} x_a & y_a & z_a & \dots & z_e & c_{11} & c_{12} & c_{13} & c_{14} & \dots & c_{34} & \bar{\bar{d}}_{1,x} & \bar{\bar{d}}_{1,y} & \bar{\bar{d}}_{1,z} & \dots & \bar{\bar{d}}_{n,z} \end{pmatrix}_{6+12+3n}^T$$



$$\Delta\Delta\Phi = \mathbf{A}\Delta\mathbf{X}, \quad \Sigma_{\Delta\Phi}$$



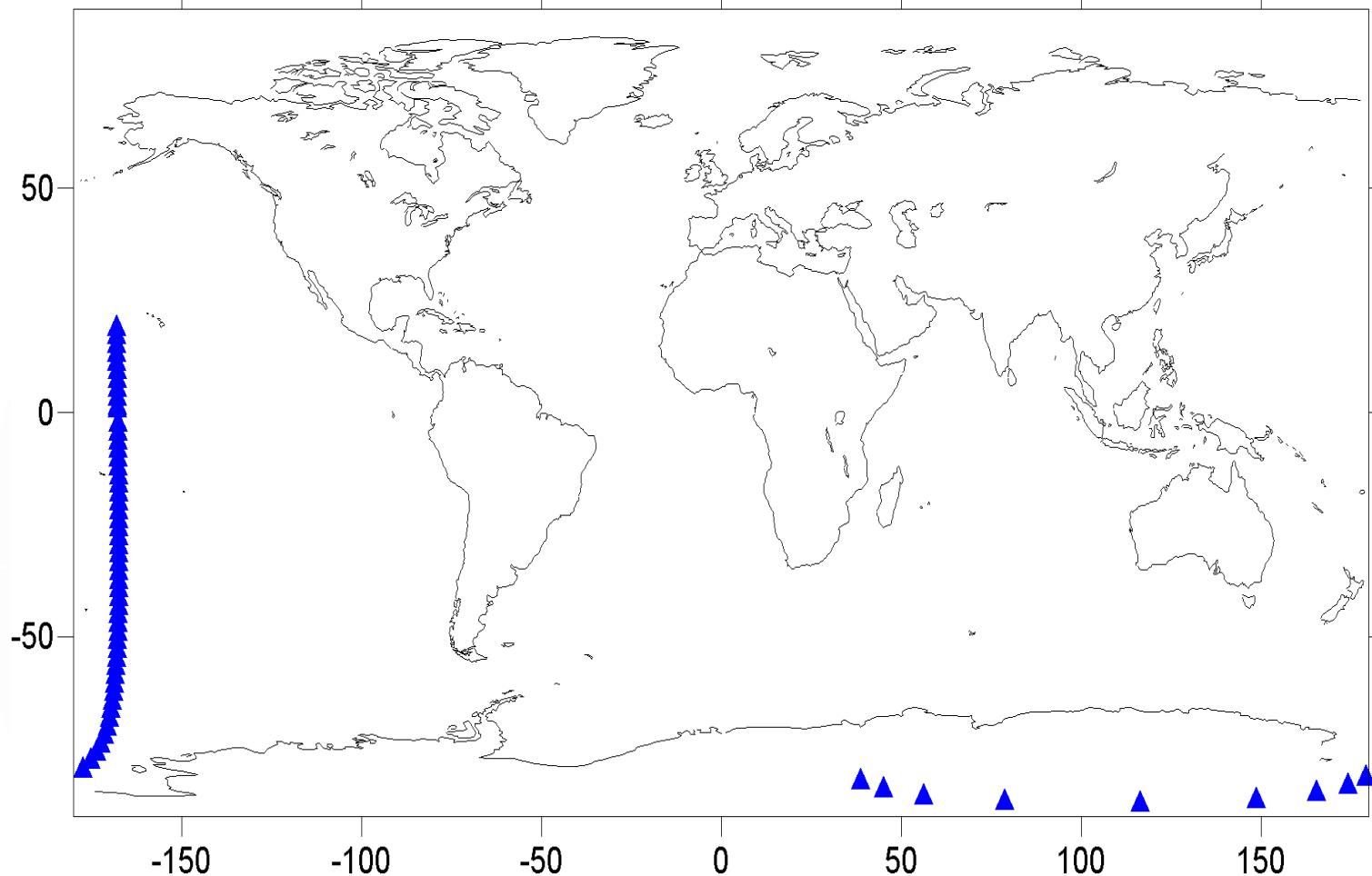
# Kinematical short arc POD sequential time differenced carrier phase



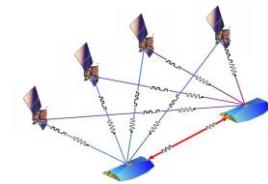
- **Unknowns:** boundary values, polynomial coefficients corrections, amplitudes of Fourier series.
- **Solutions:** Gauss-Markov model
- **Convergence & accuracy:** after a few iterations,  $\sim$  cm.



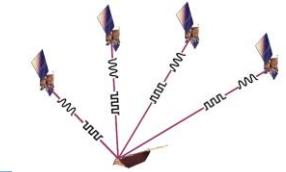
✓ LEO short arc orbit is continuous and velocities and another kinematical parameters can be derived from the estimated orbit parameters.



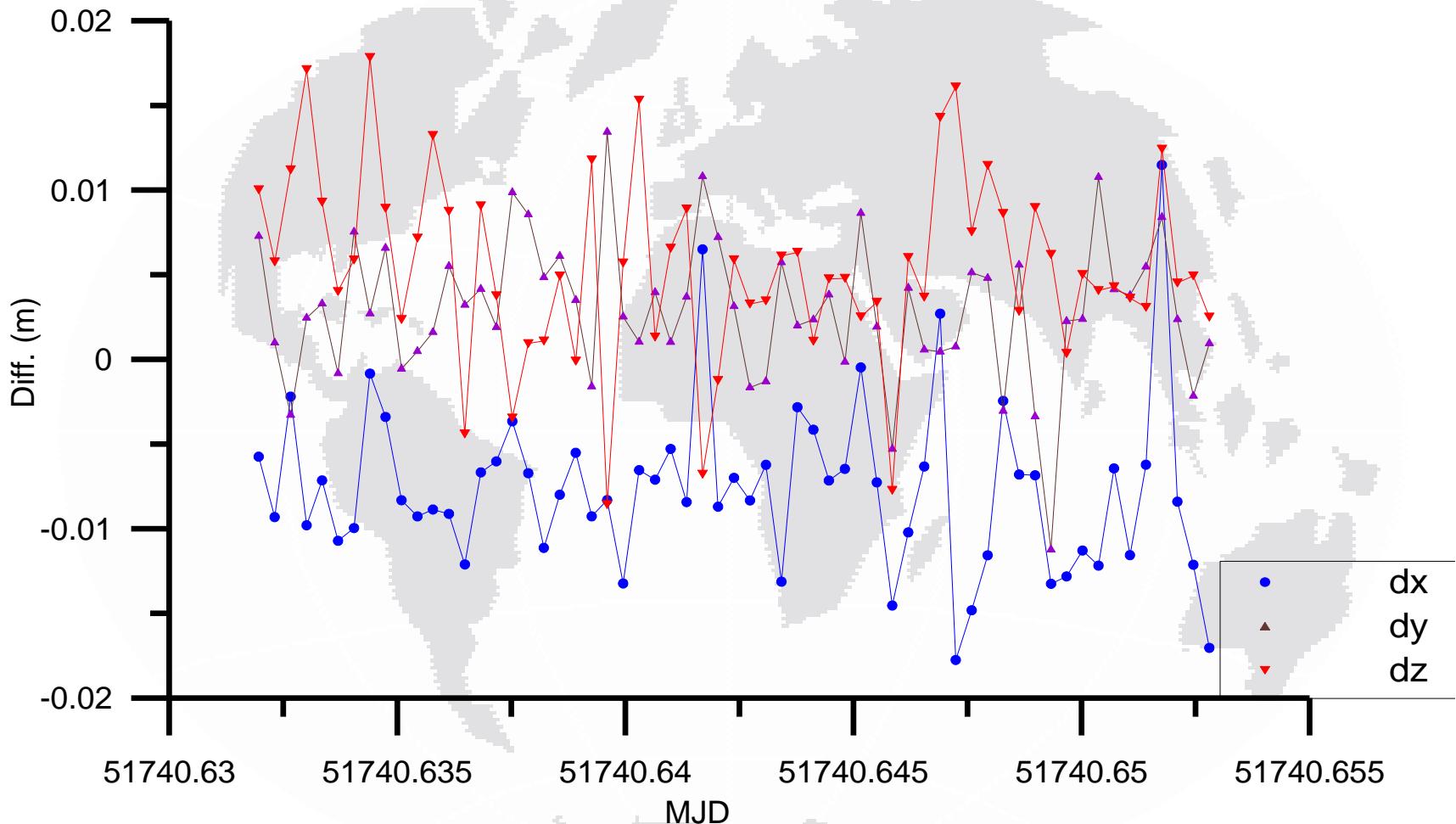
30 minutes of CHAMP satellite [ 2000 07 15 15h 10m – 15h 40m ]



# Kinematical short arc POD sequential time differenced carrier phase

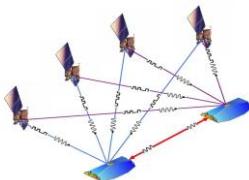


Difference plot between estimated short arc absolute positions  
with observation precision=0.01 m & given positions

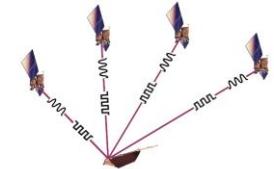




# Kinematical (short arcs) POD real case



## Carrier phase GPS SST observation



$$\Phi_{r,i}^s(t) = \left| \mathbf{R}_z(\omega_e \cdot \varepsilon_r^s) \mathbf{r}^s(t - \varepsilon_r^s) - \mathbf{r}_r(t) \right| + c \left[ dt^s(t - \varepsilon_r^s) - dt_r(t) \right] + \lambda N_r^s + I_i^r(t) + d_O^s(t) + d_R^r(t) - d_R^s(t) + d_{C,i}^r(t) + d_{V,i}^r(t) + d_{M,P_i}(t) + e_{\Phi_i}$$

$s, r$

GPS, LEO indices,

$\varepsilon_r^s$

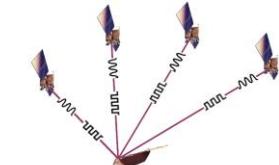
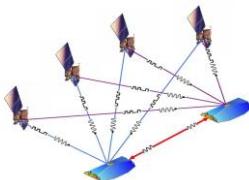
Travelling time between GPS & LEO,

$\mathbf{r}^s(t - \varepsilon_r^s), dt^s(t - \varepsilon_r^s)$

GPS position, clock offset **at sending time**,

$\mathbf{r}_r(t), dt_r(t)$

LEO position, clock offset **at receiving time**



## Carrier phase GPS SST observation...

$I_i^r(t)$

$d_O^s(t)$

$d_R^s(t), d_R^r(t)$

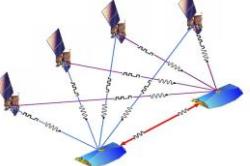
$d_{C,i}^r(t), d_{V,i}^r(t)$

$d_{M,P_i}(t)$

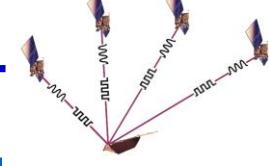
- for single frequency receiver, the IONEX model can be used to model the ionosphere error term,
- for dual frequency receiver, the ionosphere free combination can be used.

**How can the errors be  
eliminated or  
modeled in GPS LEO SST  
observations?**

multipath effect can be minimized through filtering SST observations w.r.t elevation of GPS satellites or applying the elevation weighting method or S/N filtering.



# Sequential time differenced carrier phase SST



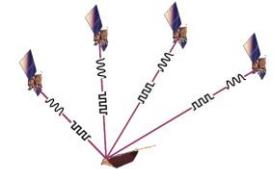
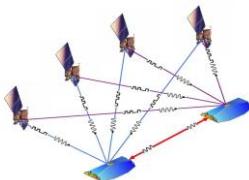
## Carrier phase ionosphere-free observation at epochs (1,2)

$$\Phi_{r,3}^s(t_1) = \left| \mathbf{R}_z(\omega_e \varepsilon_1) \mathbf{r}^s(t_1 - \varepsilon_1) - \mathbf{r}_r(t_1) \right| + \lambda_3 N_{r,3}^s + c \left[ dt^s(t_1 - \varepsilon_1) - dt_r(t_1) \right] + \\ d_O^s(t_1) + d_R^r(t_1) - d_R^s(t_1) + d_{C,3}^r(t_1) + d_{V,3}^r(t_1) + d_{M,\Phi_3}(t_1) + e_{\Phi_3}$$

$$\Phi_{r,3}^s(t_2) = \left| \mathbf{R}_z(\omega_e \varepsilon_2) \mathbf{r}^s(t_2 - \varepsilon_2) - \mathbf{r}_r(t_2) \right| + \lambda_3 N_{r,3}^s + c \left[ dt^s(t_2 - \varepsilon_2) - dt_r(t_2) \right] + \\ d_O^s(t_2) + d_R^r(t_2) - d_R^s(t_2) + d_{C,3}^r(t_2) + d_{V,3}^r(t_2) + d_{M,\Phi_3}(t_2) + e_{\Phi_3}$$

## sequential time difference ionosphere-free carrier phase observation between epochs (1,2)

$$\Delta \Phi_{r,3}^s(t_1, t_2) = \left| \mathbf{R}_z(\omega_e \varepsilon_2) \mathbf{r}^s(t_2 - \varepsilon_2) - \mathbf{r}_r(t_2) \right| - \left| \mathbf{R}_z(\omega_e \varepsilon_1) \mathbf{r}^s(t_1 - \varepsilon_1) - \mathbf{r}_r(t_1) \right| - \\ c \Delta dt_r(t_1, t_2) + e_{\Delta \Phi_3}$$



## Kinematical short arc POD

Ionosphere free sequential time differenced carrier phase observations can be written as:

$$\Delta\Phi_{r,3}^s(t_1, t_2) = \left| \mathbf{R}_z(\omega_e \varepsilon_2) \mathbf{r}^s(t_2 - \varepsilon_2) - \mathbf{r}_r(t_2) \right| - \left| \mathbf{R}_z(\omega_e \varepsilon_1) \mathbf{r}^s(t_1 - \varepsilon_1) - \mathbf{r}_r(t_1) \right| - c\Delta dt_r(t_1, t_2) + e_{\Delta\Phi_3}$$



$$\Delta\Phi_{r,3}^s(t_1, t_2) =$$

$$\left| \mathbf{R}_z(\omega_e \varepsilon_2) \mathbf{r}^s(t_2 - \varepsilon_2) - \left( \mathbf{r}_a \frac{\sin((1-\tau_2)N)}{\sin(N)} + \mathbf{r}_e \frac{\sin(\tau_2 N)}{\sin(N)} + \mathbf{C}_{3 \times 4} \mathbf{P}(\tau_2) + \sum_{f=1}^n \bar{\mathbf{d}}_f \sin(\pi f \tau_2) \right) \right| -$$

$$\left| \mathbf{R}_z(\omega_e \varepsilon_1) \mathbf{r}^s(t_1 - \varepsilon_1) - \left( \mathbf{r}_a \frac{\sin((1-\tau_1)N)}{\sin(N)} + \mathbf{r}_e \frac{\sin(\tau_1 N)}{\sin(N)} + \mathbf{C}_{3 \times 4} \mathbf{P}(\tau_1) + \sum_{f=1}^n \bar{\mathbf{d}}_f \sin(\pi f \tau_1) \right) \right| -$$

$$c\Delta dt_r(t_1, t_2) + e_{\Delta\Phi_3}$$

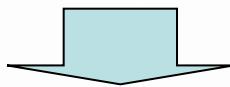
# Gauss-Markov model

The Gauss-Markov model for one epoch,

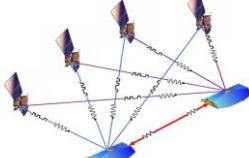
$$\begin{pmatrix} \Delta\Delta\Phi_r^{s_1}(t_1, t_2) \\ \vdots \\ \Delta\Delta\Phi_r^{s_j}(t_1, t_2) \\ \vdots \\ \Delta\Delta\Phi_r^{s_{m_1}}(t_1, t_2) \end{pmatrix} = \begin{pmatrix} \mathbf{A}_{\mathbf{r}_r}^{s_1}(t_2)\mathbf{A}_{\mathbf{X}_r}^{\mathbf{r}_r}(t_2) - \mathbf{A}_{\mathbf{r}_r}^{s_1}(t_1)\mathbf{A}_{\mathbf{X}_r}^{\mathbf{r}_r}(t_1) \\ \vdots \\ \mathbf{A}_{\mathbf{r}_r}^{s_j}(t_2)\mathbf{A}_{\mathbf{X}_r}^{\mathbf{r}_r}(t_2) - \mathbf{A}_{\mathbf{r}_r}^{s_j}(t_1)\mathbf{A}_{\mathbf{X}_r}^{\mathbf{r}_r}(t_1) \\ \vdots \\ \mathbf{A}_{\mathbf{r}_r}^{s_{m_1}}(t_2)\mathbf{A}_{\mathbf{X}_r}^{\mathbf{r}_r}(t_2) - \mathbf{A}_{\mathbf{r}_r}^{s_{m_1}}(t_1)\mathbf{A}_{\mathbf{X}_r}^{\mathbf{r}_r}(t_1) \end{pmatrix} [\mathbf{X}_r - \mathbf{X}_r^0] + \begin{pmatrix} 1 & \dots & 0 \\ \vdots & \ddots & \vdots \\ 1 & 0 & 0 \\ \vdots & \ddots & \vdots \\ 1 & \dots & 0 \end{pmatrix} [\mathbf{X}_{\Delta t} - \mathbf{X}_{\Delta t}^0]$$

$$\mathbf{X}_r = \begin{pmatrix} x_a & y_a & z_a & \dots & z_e & c_{11} & c_{12} & c_{13} & c_{14} & \dots & c_{34} & \bar{\bar{d}}_{1,x} & \bar{\bar{d}}_{1,y} & \bar{\bar{d}}_{1,z} & \dots & \bar{\bar{d}}_{n,z} \end{pmatrix}_{6+12+3n}^T$$

$$\mathbf{X}_{\Delta t} = (\Delta cdt_{12} \quad \dots \quad \Delta cdt_{(m-1)m})_{m-1}^T$$



$$\Delta\Delta\Phi = \mathbf{A}\Delta\mathbf{X}, \quad \Sigma_{\Delta\Phi}$$



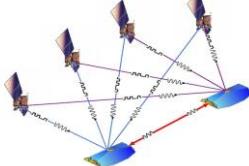
# Kinematical short arc POD sequential time differenced carrier phase

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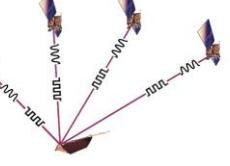
- **Initial value:** unknowns initial values can be derived from code estimated positions at the first step,
- **Unknowns:** boundary values, polynomial coefficients corrections, amplitudes of Fourier series,
- **Solutions:** Gauss-Markov model,
- **Convergence & accuracy:** after  $\sim$ a few iterations,  $\sim$  cm.



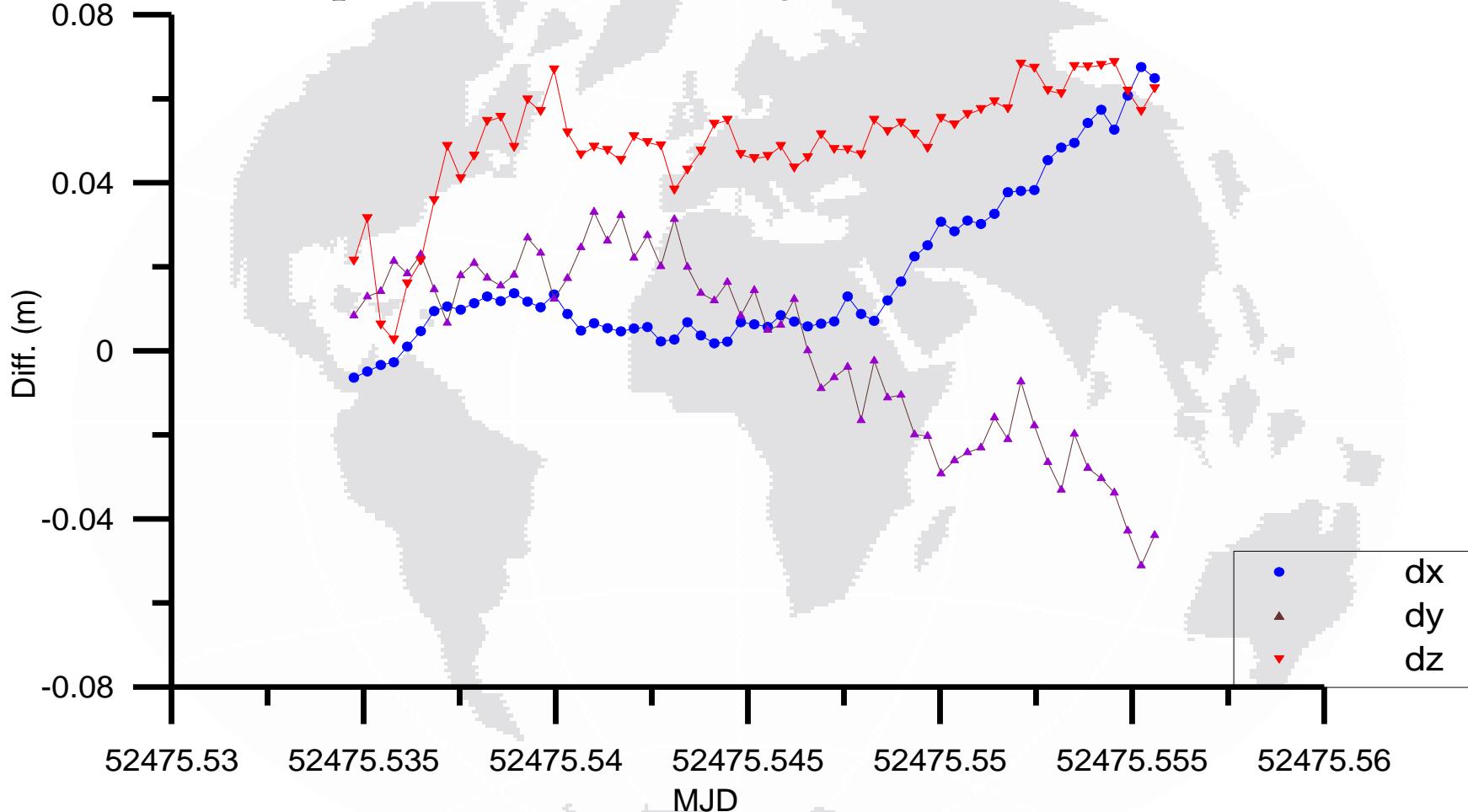
- ✓ LEO short arc orbit is continuous and velocities and another kinematical parameters can be derived from the estimated LEO short arc parameters.

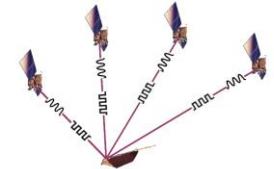
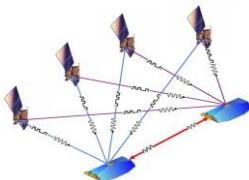


# Kinematical short arc POD with sequential time differenced carrier phase



Difference plot between absolute positions with carrier phase observation precision = 0.01 m & given GFZ CHAMP PSO orbit



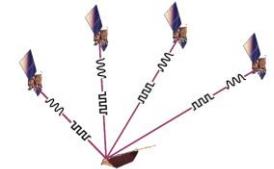
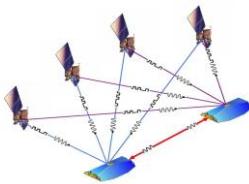


## Conclusions & remarks (kinematical)

- From sequential differenced carrier phase SST observations, LEO absolute positions and clock offsets can be estimated at every epoch with a sufficient (not essential 4) number of GPS satellites and good satellite geometry,
- An accuracy of **cm** can be expected for the sequential time diffenced carrier phase SST data processing, but **DOP!** isn't crucial,
- The resulting LEO orbit is given **continuous** (without gaps) & smoother than geometrical POD,
- Kinematical LEO orbit can be used to recover the Earth's gravity field with the POD recovery concept. (**kinematical parameters can be derived analytically**).



# Dynamical (short arcs) POD simulation case



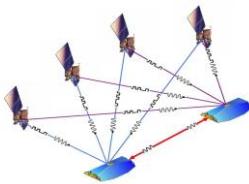
## Dynamical short arc POD

$$\mathbf{r}(t) = \mathbf{r}_a + \frac{\sin((1-\tau)N)}{\sin(N)} + \mathbf{r}_e + \frac{\sin(\tau N)}{\sin(N)} + \mathbf{C}^T \mathbf{P}(\tau) + \sum_{v=1}^n \bar{\mathbf{d}}_v \sin(v\pi\tau)$$

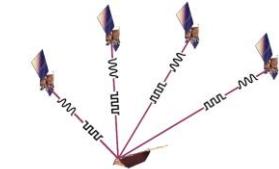
Boundary positions have been estimated by a  
Polynomial & Fornieles series coefficients have  
been determined from given Earth gravity

**field  
disadvantage:**

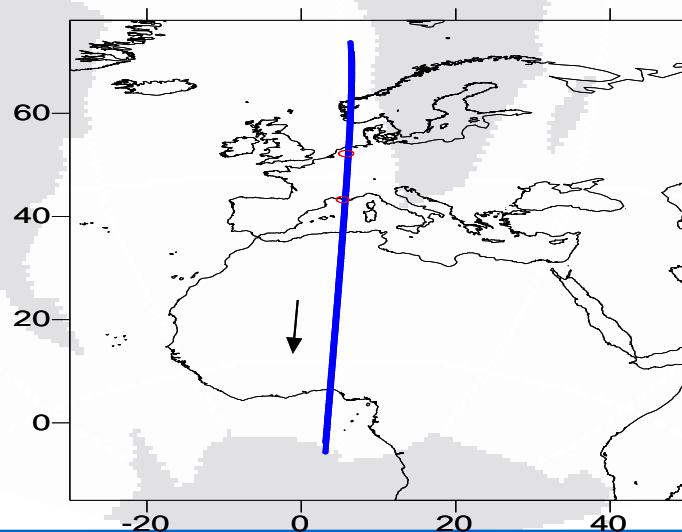
the dynamical orbits can't be used directly to recover the Earth gravity field, but estimated dynamical parameters can be used as initial values to kinematical POD procedure.

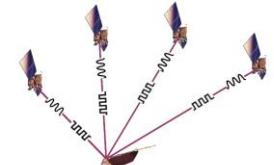
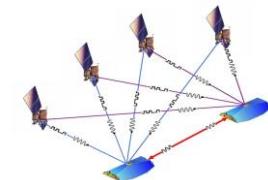


## Dynamical short arc POD-simulation

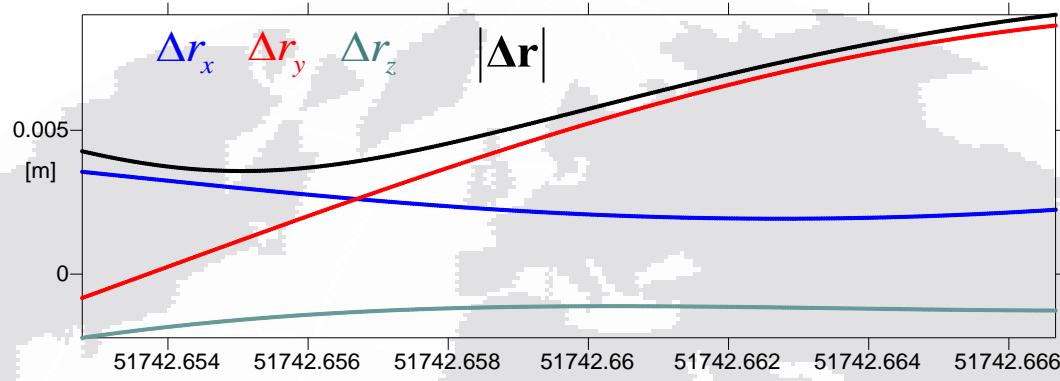


- Short arcs of GRACE twin satellites above Europe and Africa have been selected,
- Observations type:
  - GPS high-low SST pseudo-range observations,
  - Earth gravity field (EGM96),
- Unknowns :
  - Boundary positions at begin & end of short arc.

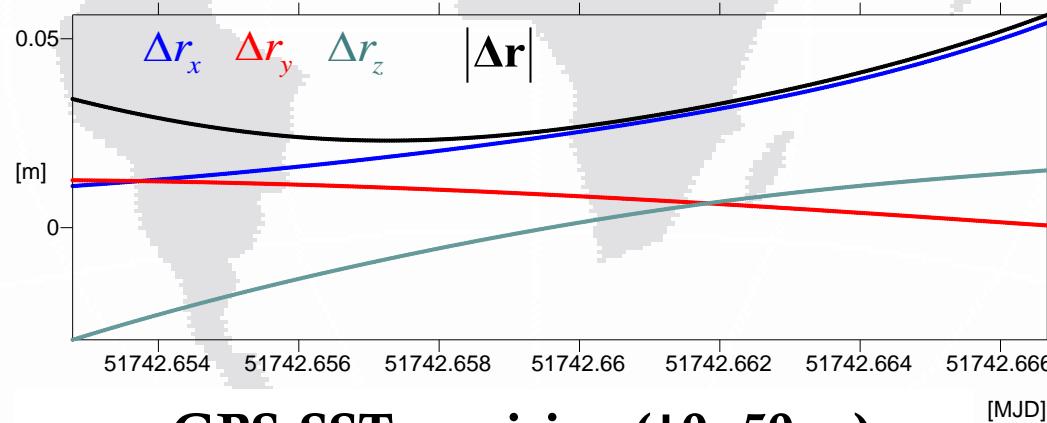




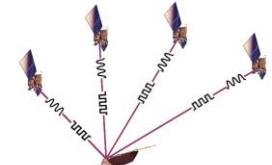
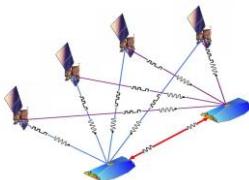
# Dynamical short arc POD-simulation results



GPS-SST precision ( $\pm 0,05$  m)



GPS-SST precision ( $\pm 0, 50$  m)

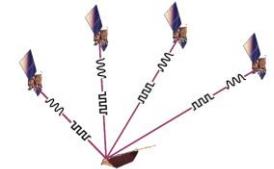
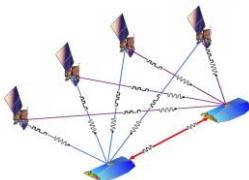


## Conclusions & remarks (dynamical)

- Dynamical LEO POD can not be used for recovery of the Earth with the POD recovery concept, but estimated parameters & LEO POD can be used as initial values for further processes as kinematical POD, etc.,,
- From dynamical method, LEO positions and clock offsets can be estimated at every epoch with a sufficient enough number (not essential 4) of GPS satellites and good satellite geometry,
- The resulting LEO orbit is given **continuous** (without gaps) and smoother than kinematical POD,



# Reduced kinematical (short arcs) POD simulation case

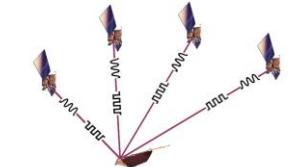
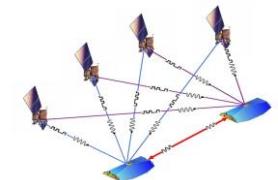


## Reduced kinematical POD

$$\mathbf{r}(t) = \mathbf{r}_a \cdot \frac{\sin((1-\tau)N)}{\sin(N)} + \mathbf{r}_e \cdot \frac{\sin(\tau N)}{\sin(N)} + \mathbf{C}^T \mathbf{P}(\tau) + \sum_{v=1}^n \bar{\mathbf{d}}_v \sin(v\pi\tau)$$

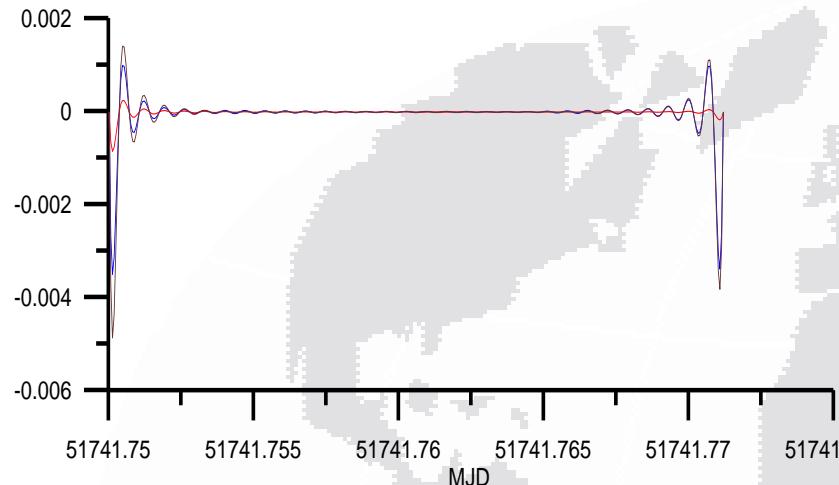
Calculated from low degree  
Earth gravity field

- Sinsus analysis to difference between real orbit and dynamical low degree orbit,
- A little dynamical information have been used in the reduced kinematical orbit determination,

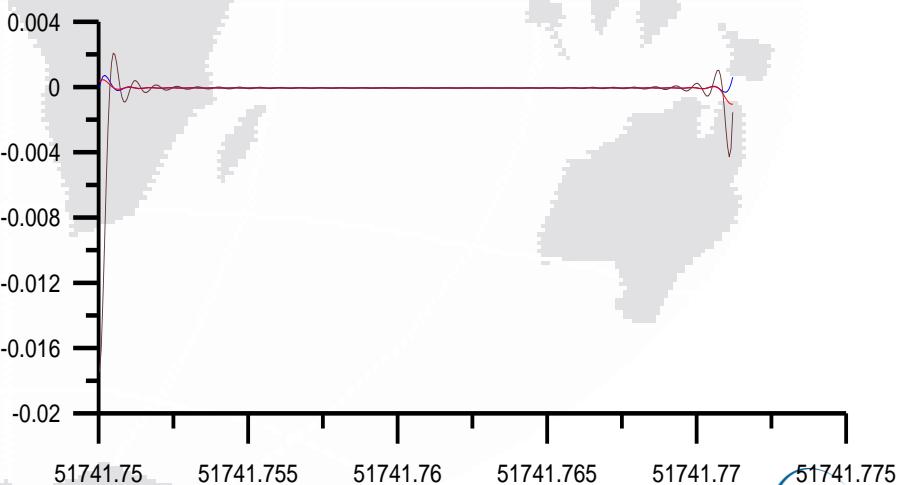
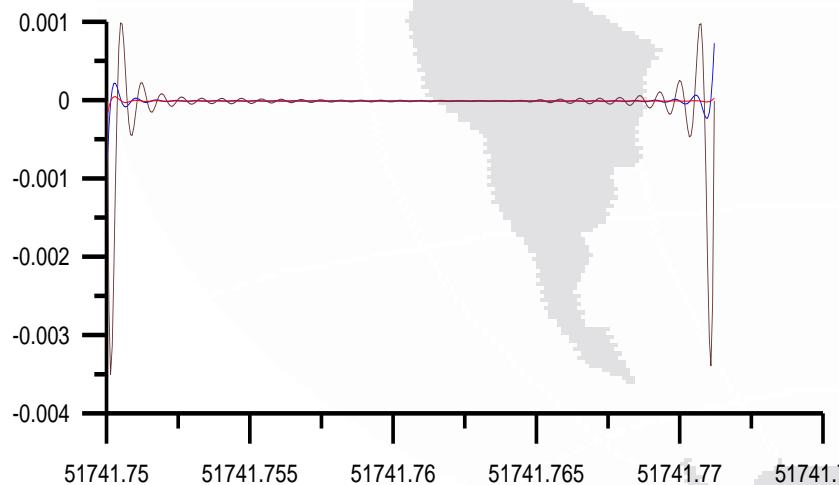
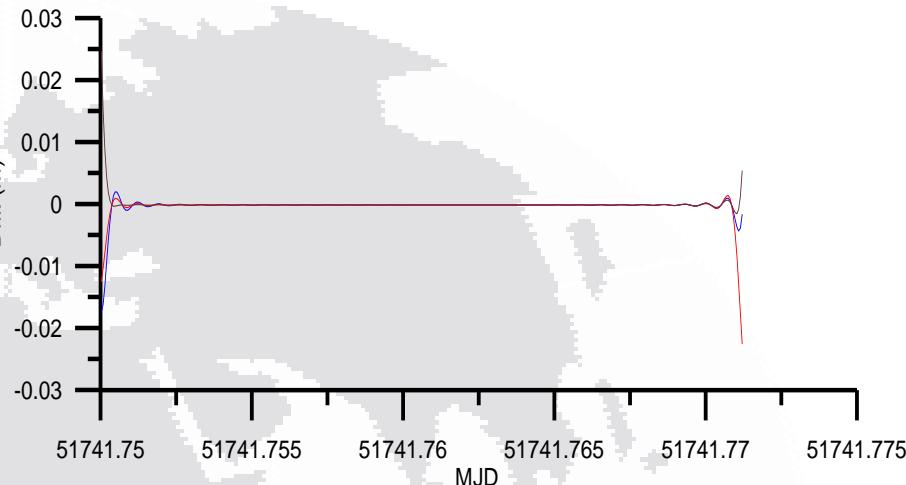


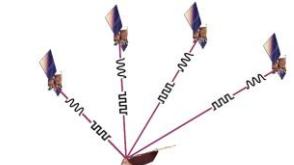
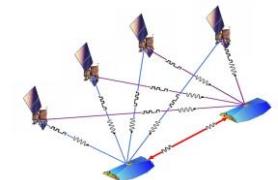
## Reduced kinematical POD, results

Earth gravity field deg. 02, error free



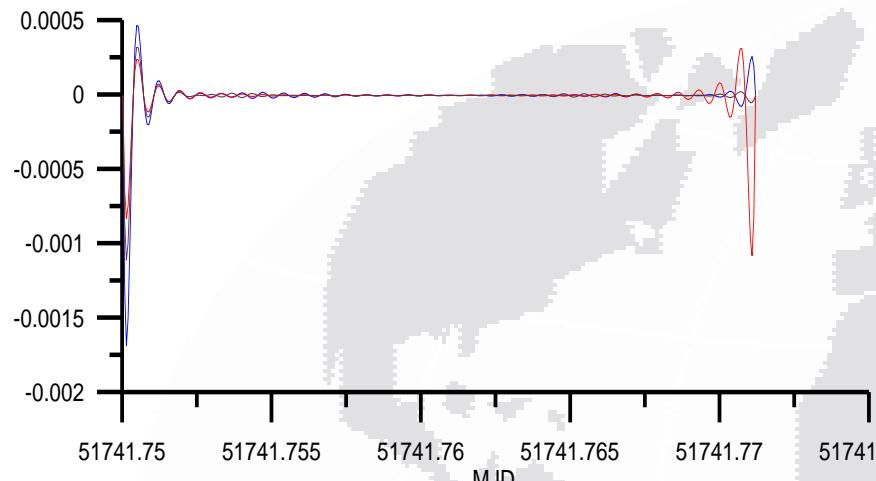
Earth gravity field deg. 02, error 05 cm



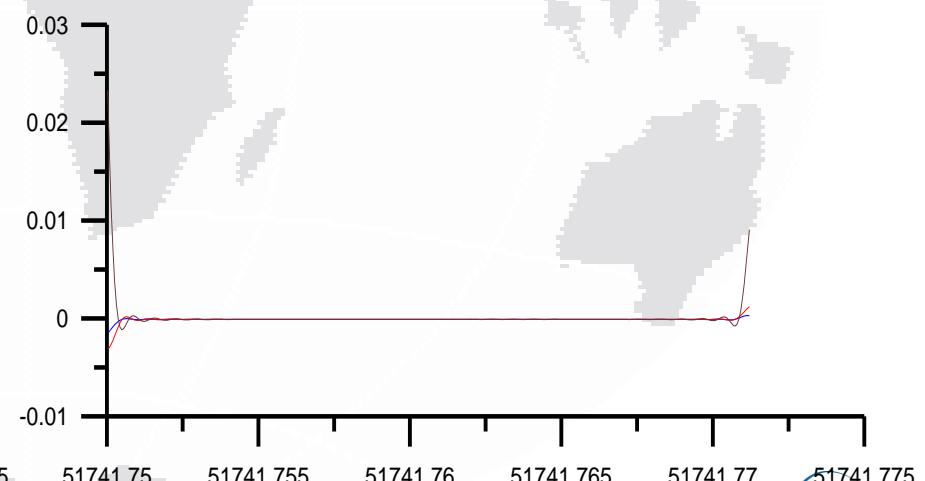
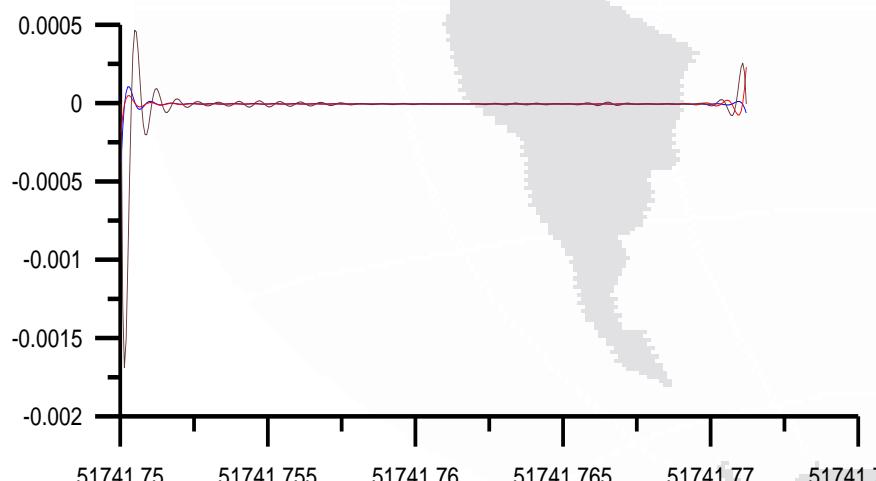
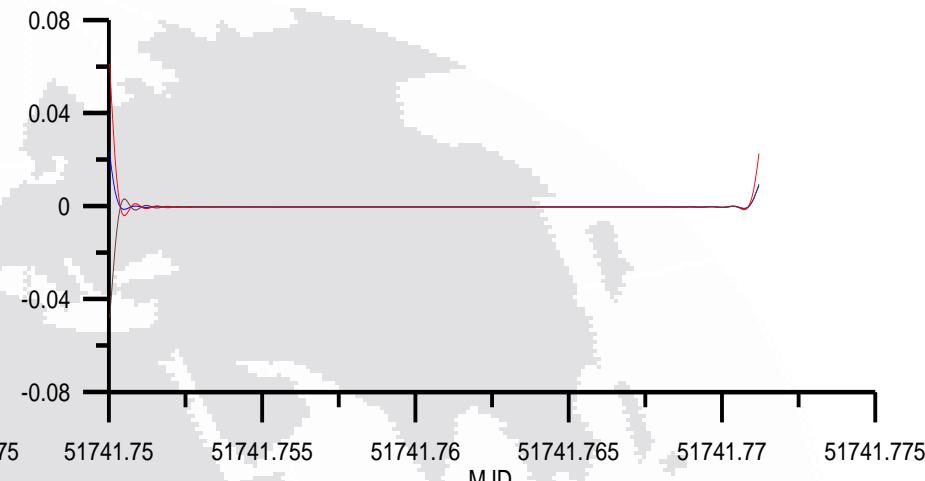


# Reduced kinematical POD, results

Earth gravity field deg. 06, error free



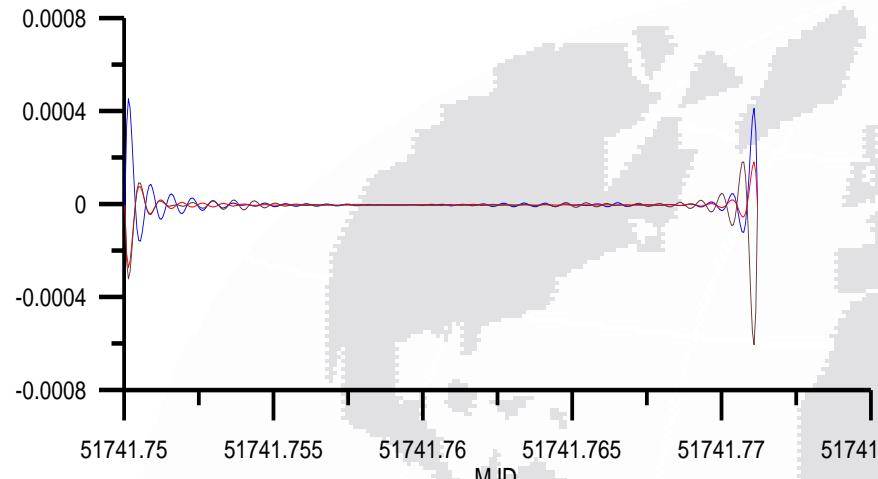
Earth gravity field deg. 06, error 05 cm



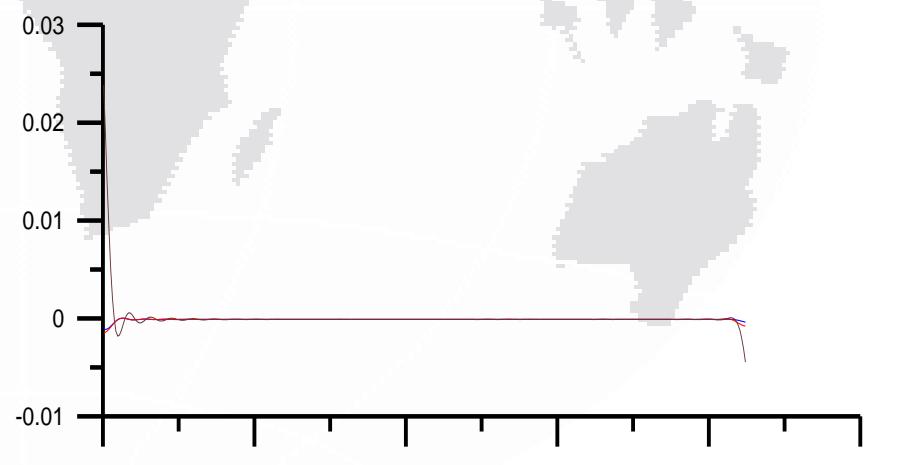
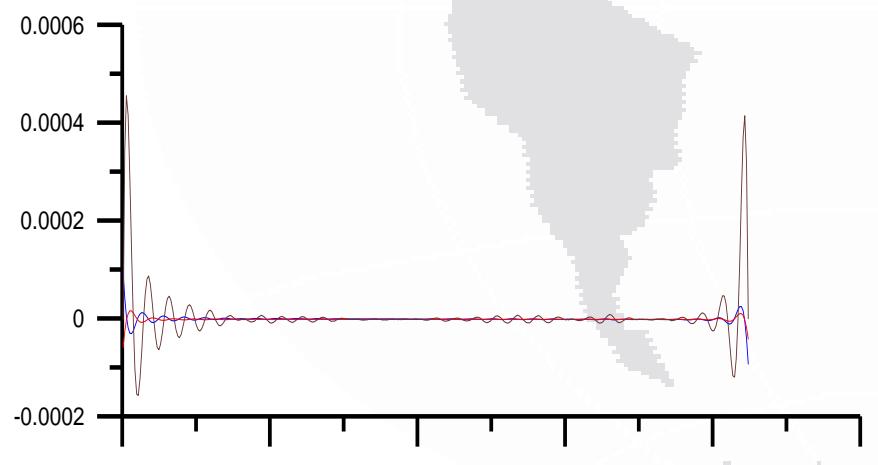
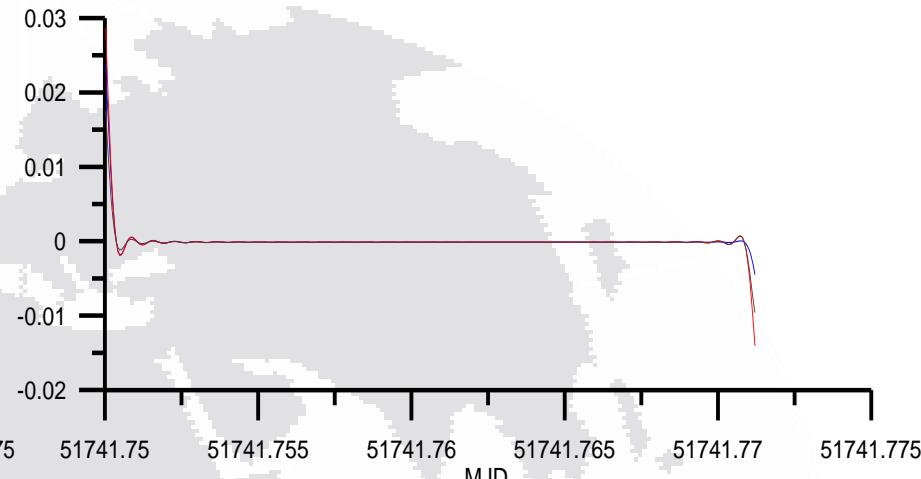


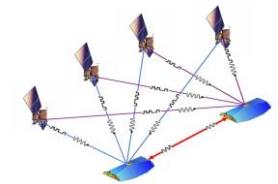
## Reduced kinematical POD, results

Earth gravity field deg. 12, error free

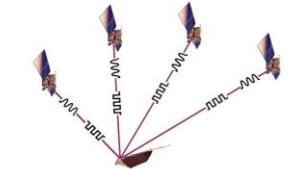


Earth gravity field deg. 12, error 05 cm

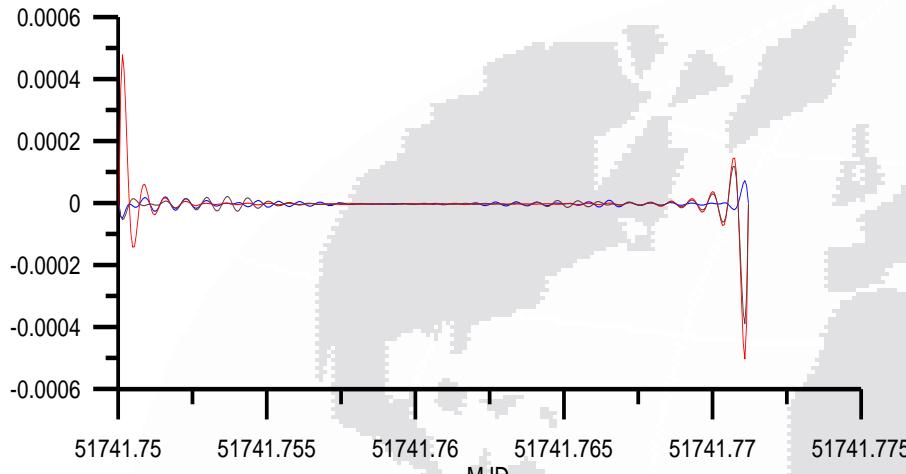




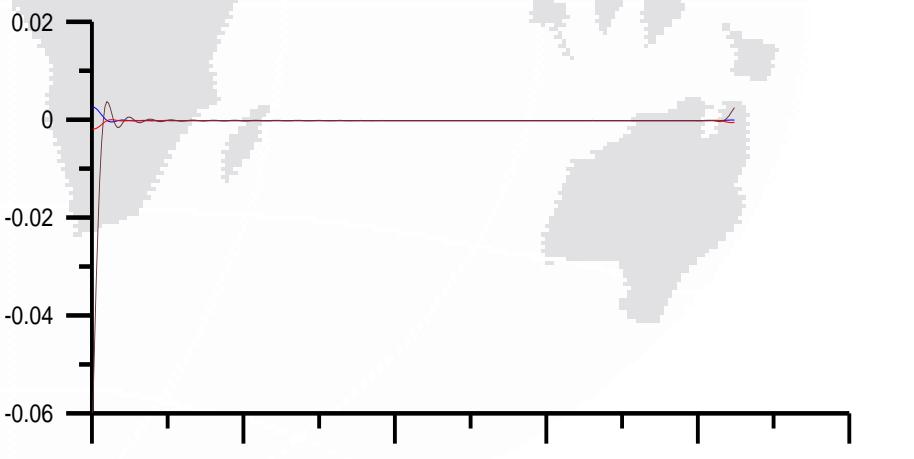
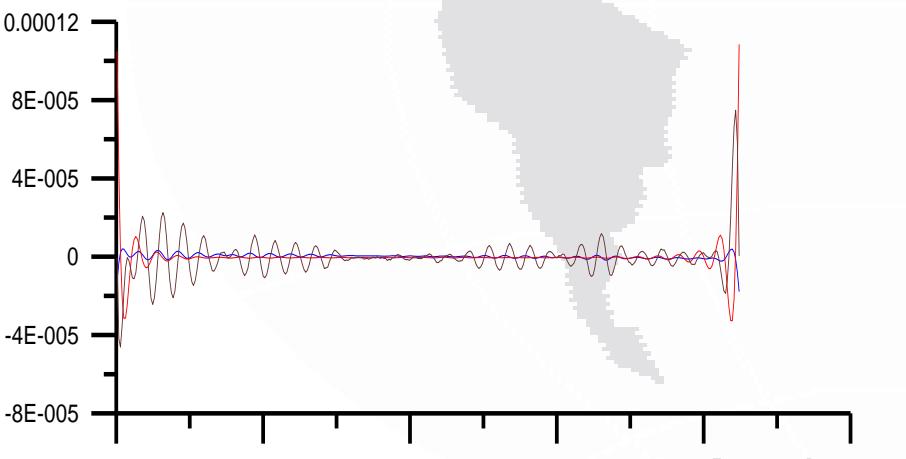
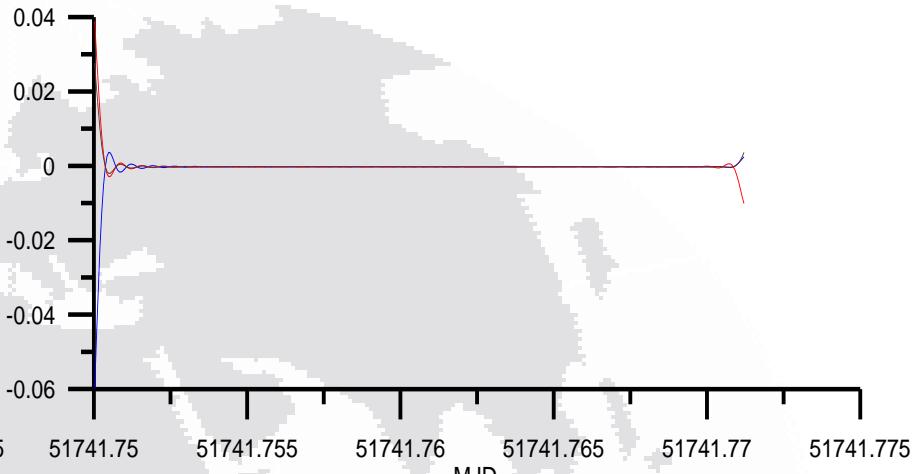
# Reduced kinematical POD, results



Earth gravity field deg. 16, error free



Earth gravity field deg. 16, error 05 cm





Thank you for your  
attentions

