Physics Letters B 784 (2018) 137-141

Contents lists available at ScienceDirect

Physics Letters B

www.elsevier.com/locate/physletb

\mathcal{N} -extended supersymmetric Calogero models

Sergey Krivonos^{a,b}, Olaf Lechtenfeld^{c,*}, Anton Sutulin^a

^a Bogoliubov Laboratory of Theoretical Physics, JINR, 141980 Dubna, Russia

^b St. Petersburg Department of V.A. Steklov Institute of Mathematics of the Russian Academy of Sciences, 27 Fontanka, St. Petersburg, Russia

^c Institut für Theoretische Physik and Riemann Center for Geometry and Physics, Leibniz Universität Hannover, Appelstrasse 2, D-30167 Hannover, Germany

ARTICLE INFO

Article history: Received 17 May 2018 Received in revised form 4 July 2018 Accepted 20 July 2018 Available online 24 July 2018 Editor: M. Cvetič

ABSTRACT

We propose a new N-extended supersymmetric su(n) spin-Calogero model. Employing a generalized Hamiltonian reduction adopted to the supersymmetric case, we explicitly construct a novel rational *n*-particle Calogero model with an arbitrary even number of supersymmetries. It features Nn^2 rather than Nn fermionic coordinates and increasingly high fermionic powers in the supercharges and the Hamiltonian.

© 2018 The Authors. Published by Elsevier B.V. This is an open access article under the CC BY license (http://creativecommons.org/licenses/by/4.0/). Funded by SCOAP³.

1. Introduction

The original rational Calogero model of *n* interacting identical particles on a line [1], pertaining to the roots of $A_1 \oplus A_{n-1}$ and given by the classical Hamiltonian

$$H = \frac{1}{2} \sum_{i=1}^{n} p_i^2 + \frac{1}{2} \sum_{i \neq j} \frac{g^2}{\left(x^i - x^j\right)^2} , \qquad (1.1)$$

has often been the subject of "supersymmetrization". In this endeavor, extended supersymmetry has turned out to be surprisingly rich. After the straightforward formulation of $\mathcal{N}=2$ supersymmetric Calogero models by Freedman and Mende [2], a barrier was encountered at $\mathcal{N} = 4$ [3]. An important step forward then was the explicit construction of the supercharges and the Hamiltonian for the $\mathcal{N}=4$ supersymmetric three-particle Calogero model [4,5], which introduced a second prepotential F besides the familiar prepotential U. However, it was found that quantum corrections modify the potential in (1.1), and that F is subject to intricate nonlinear differential equations, the WDVV equations, beyond the three-particle case. These results were then confirmed and elucidated in a superspace description [6]. Finally, extending the system by a single harmonic degree of freedom (su(2) spin variables [7])it was possible to write down a unique osp(4|2) symmetric fourparticle Calogero model [8].¹ A detailed discussion concerning the

* Corresponding author.

lechtenf@itp.uni-hannover.de (O. Lechtenfeld), sutulin@theor.jinr.ru (A. Sutulin).
 ¹ Here and in the above history, the goal is a bosonic potential exactly as in (1.1).
 Models with more general interactions can be found for any number of particles.

supersymmetrization of the Calogero models can be found in the review [9].

It seems that a guiding principle was missing for the construction of extended supersymmetric Calogero models. Indeed, while for $n \leq 3$ translation and (super-)conformal symmetry almost completely defines the system, the $n \geq 4$ cases admit a lot of freedom which cannot *a priori* be fixed. In the bosonic case, such a guiding principle exists [10]. The Calogero model as well as its different extensions (see, e.g. [11–13]) are closely related with matrix models and can be obtained from them by a reduction procedure (see [14] for first results and [15] for a review). If we want to employ this principle also for finding extended supersymmetric Calogero models, then the two main steps are

- supersymmetrization of a matrix model
- supersymmetrization of the reduction procedure or proper gauge fixing.

This idea is not new. It has successfully been employed in [16–19]. The resulting supersymmetric systems feature

- a large number of fermions far more than the 4*n* fermions expected in an $\mathcal{N}=4$ *n*-particle system within the standard (but unsuccessful!) approach
- a rather complicated structure of the supercharges and the Hamiltonian, with fermionic polynomials of maximal degree
- a variety of bosonic potentials, including *su*(2) *spin*-Calogero interactions

https://doi.org/10.1016/j.physletb.2018.07.036





E-mail addresses: krivonos@theor.jinr.ru (S. Krivonos),

^{0370-2693/© 2018} The Authors. Published by Elsevier B.V. This is an open access article under the CC BY license (http://creativecommons.org/licenses/by/4.0/). Funded by SCOAP³.

but they do not contain a genuine $\mathcal{N}=4$ supersymmetric Calogero model, i.e. one with a mere pairwise inverse-square no-spin bosonic potential.

Here we use the same guiding principle and start with the bosonic su(n) spin-Calogero model in the Hamiltonian approach. We then provide an \mathcal{N} -extended supersymmetrization of this system. It is important that we do *not a priori* fix a realization for the su(n) generators. Finally we generalize the reduction procedure to the \mathcal{N} -extended system and find the first \mathcal{N} -extended supersymmetric Calogero model, for *any* even number of supersymmetries.

2. N-extended supersymmetric Calogero model

2.1. Bosonic Calogero model from hermitian matrices

It is well known that the rational *n*-particle Calogero model [1] can be obtained by Hamiltonian reduction from the hermitian matrix model [10,14]. Adapted to our purposes, the procedure reads as follows. One starts from the su(n) spin generalization [12] of the standard Calogero model, as given by

$$H = \frac{1}{2} \sum_{i=1}^{n} p_i^2 + \frac{1}{2} \sum_{i \neq j}^{n} \frac{\ell_{ij} \ell_{ji}}{\left(x^i - x^j\right)^2} \,. \tag{2.1}$$

The particles are described by their coordinates x^i and momenta p_i together with their internal degrees of freedom encoded in the angular momenta $(\ell_{ij})^{\dagger} = \ell_{ji}$ with $\sum_i \ell_{ii} = 0$. The non-vanishing Poisson brackets are

$$\left\{x^{i}, p_{j}\right\} = \delta_{j}^{i} \quad \text{and} \quad \left\{\ell_{ij}, \ell_{km}\right\} = i\left(\delta_{im}\ell_{kj} - \delta_{kj}\ell_{im}\right) . \quad (2.2)$$

The Hamiltonian (2.1) follows directly from the free hermitian matrix model (for details see [15]).

To get the standard Calogero Hamiltonian (1.1) from (2.1) one has to reduce the angular sector of the latter, in two steps. Firstly, one (weakly) imposes the constraints

$$\ell_{11} \approx \ell_{22} \approx \ldots \approx \ell_{nn} \approx 0.$$
(2.3)

They commute with the Hamiltonian (2.1) and with each other, hence are of first class. To resolve them one introduces auxiliary complex variables v_i and $\bar{v}_i = (v_i)^{\dagger}$ obeying the Poisson brackets

$$\left\{\nu_i, \bar{\nu}_j\right\} = -i\delta_{ij} \tag{2.4}$$

and realizes the su(n) generators ℓ_{ij} as

$$\hat{\ell}_{ij} = -\nu_i \bar{\nu}_j + \frac{1}{n} \delta_{ij} \sum_k^n \nu_k \bar{\nu}_k \,. \tag{2.5}$$

Secondly, passing to polar variables r_i and ϕ_i defined as

$$v_i = r_i e^{i\phi_i}$$
 and $\bar{v}_i = r_i e^{-i\phi_i}$ \Rightarrow $\{r_i, \phi_j\} = \frac{1}{2r_i}\delta_{ij}$,
(2.6)

the constraints (2.3) are resolved by putting

 $r_1 \approx r_2 \approx \ldots \approx r_n$ (2.7)

Plugging this solution into the Hamiltonian (2.1) one may additionally fix n-1 angles ϕ_i , say

$$\phi_1 \approx \phi_2 \approx \ldots \approx \phi_{n-1} \approx 0 . \tag{2.8}$$

At this stage the 2n variables $\{r_i, \phi_i\}$ are reduced to the two variables r_n and ϕ_n . However, the reduced Hamiltonian does not depend on ϕ_n and has the form

$$H_{\rm red} = \frac{1}{2} \sum_{i=1}^{n} p_i^2 + \frac{1}{2} \sum_{i \neq j}^{n} \frac{r_n^4}{\left(x^i - x^j\right)^2} \,. \tag{2.9}$$

Therefore

$$\{H_{\rm red}, r_n\} \approx 0$$
 and $r_n^2 \approx {\rm const} =: g$, (2.10)

and the reduced Hamiltonian H_{red} coincides with the standard *n*-particle rational Calogero Hamiltonian. We note that in the bosonic case most reduction steps are not needed, because the Hamiltonian (2.1) does not depend on the angles ϕ_i at all. However, in the supersymmetric case all reduction steps will be important.

In what follows we will construct an N-extended supersymmetric generalization of the Hamiltonian (2.1) and perform the supersymmetric version of the reduction just discussed, finishing with an N-extended supersymmetric Calogero model, for N = 2M and M = 1, 2, 3, ...

2.2. N-extended supersymmetric su(n) spin-Calogero model

On the outset we have to clarify what is the minimal number of fermionic variables necessary to realize an $\mathcal{N} = 2M$ supersymmetric extension of the su(n) spin-Calogero model (2.1). Clearly, as partners to the bosonic coordinates x^i one needs $\mathcal{N}n$ fermions ψ_i^a and $\bar{\psi}_{ia}$ with $a = 1, 2, \dots M$. However, this is not enough to construct \mathcal{N} supercharges Q^a and \overline{Q}_b which must generate the $\mathcal{N} = 2M$ superalgebra

$$\left\{Q^{a}, \overline{Q}_{b}\right\} = -2i \,\delta^{a}_{b} \, H \qquad \text{and} \qquad \left\{Q^{a}, Q^{b}\right\} = \left\{\overline{Q}_{a}, \overline{Q}_{b}\right\} = 0 \,.$$
(2.11)

The reason is simple: to generate the potential term $\sum_{i\neq j}^{n} \frac{\ell_{ij}\ell_{ji}}{(x^i-x^j)^2}$ in the Hamiltonian, the supercharges Q^a and \overline{Q}_b must contain the terms

$$i\sum_{i\neq j}^{n} \frac{\ell_{ij}\rho_{ji}^{a}}{x^{i}-x^{j}} \quad \text{and} \quad -i\sum_{i\neq j}^{n} \frac{\ell_{ji}\bar{\rho}_{ija}}{x^{i}-x^{j}}, \quad (2.12)$$

respectively, where ρ_{ij}^a and $\bar{\rho}_{ija}$ are some additional fermionic variables. These fermions cannot be constructed from ψ_i^a or $\bar{\psi}_{ia}$. Hence, we are forced to introduce $\mathcal{N}n(n-1)$ further independent fermions ρ_{ij}^a and $\bar{\rho}_{ija}$ subject to $\rho_{ii}^a = \bar{\rho}_{iia} = 0$ for each value of the index *i*. In total, we thus utilize $\mathcal{N}n^2$ fermions of type ψ or ρ , which we demand to obey the following Poisson brackets,

$$\{\psi_i^a, \bar{\psi}_{jb}\} = -i \delta_b^a \delta_{ij}, \quad \{\rho_{ij}^a, \bar{\rho}_{kmb}\} = -i \delta_b^a \delta_{im} \delta_{jk},$$

$$\text{with} \quad (\rho_{ij}^a)^\dagger = \bar{\rho}_{jia} \quad \text{and} \quad \rho_{ii}^a = \bar{\rho}_{iia} = 0.$$

$$(2.13)$$

The next important ingredient of our construction is the composite object

$$\Pi_{ij} = \sum_{a=1}^{M} \left[\left(\psi_{i}^{a} - \psi_{j}^{a} \right) \bar{\rho}_{ija} + \left(\bar{\psi}_{ia} - \bar{\psi}_{ja} \right) \rho_{ij}^{a} + \sum_{k=1}^{n} \left(\rho_{ik}^{a} \bar{\rho}_{kja} + \bar{\rho}_{ika} \rho_{kj}^{a} \right) \right] \quad \Rightarrow \quad (\Pi_{ij})^{\dagger} = \Pi_{ji} .$$
(2.14)

One may check that, with respect to the brackets (2.13), the Π_{ij} form an su(n) algebra just like the ℓ_{ii} ,

$$\left\{\Pi_{ij}, \Pi_{km}\right\} = i \left(\delta_{im} \Pi_{kj} - \delta_{kj} \Pi_{im}\right) , \qquad (2.15)$$

and they Poisson-commute with the our fermions as follows,

$$\{\Pi_{ij}, \psi_k^a\} = \mathbf{i} \left(\delta_{ik} - \delta_{jk}\right) \rho_{ij}^a , \{\Pi_{ij}, \rho_{km}^a\} = -\mathbf{i} \,\delta_{im} \delta_{jk} \left(\psi_i^a - \psi_j^a\right) - \mathbf{i} \delta_{jk} \rho_{im}^a + \mathbf{i} \delta_{im} \rho_{kj}^a , \{\Pi_{ij}, \bar{\psi}_{ka}\} = \mathbf{i} \left(\delta_{ik} - \delta_{jk}\right) \bar{\rho}_{ija} , \{\Pi_{ij}, \bar{\rho}_{kma}\} = -\mathbf{i} \,\delta_{im} \delta_{jk} \left(\bar{\psi}_{ia} - \bar{\psi}_{ja}\right) - \mathbf{i} \delta_{jk} \bar{\rho}_{ima} + \mathbf{i} \delta_{im} \bar{\rho}_{kja} .$$

$$(2.16)$$

It is a matter of straightforward calculation to check that the supercharges

$$Q^{a} = \sum_{i=1}^{n} p_{i}\psi_{i}^{a} + i\sum_{i\neq j}^{n} \frac{(\ell_{ij} + \Pi_{ij})\rho_{ji}^{a}}{x^{i} - x^{j}} \quad \text{and} \\ \overline{Q}_{b} = \sum_{i=1}^{n} p_{i}\bar{\psi}_{ib} - i\sum_{i\neq j}^{n} \frac{\bar{\rho}_{ijb}(\ell_{ji} + \Pi_{ji})}{x^{i} - x^{j}}$$
(2.17)

obey the N = 2M superalgebra (2.11) with the Hamiltonian

$$H = \frac{1}{2} \sum_{i=1}^{n} p_i^2 + \frac{1}{2} \sum_{i \neq j}^{n} \frac{\left(\ell_{ij} + \Pi_{ij}\right) \left(\ell_{ji} + \Pi_{ji}\right)}{\left(x^i - x^j\right)^2} , \qquad (2.18)$$

modulo the first-class constraints²

$$\chi_i := \ell_{ii} + \Pi_{ii} \approx 0 \quad \forall i , \qquad (2.19)$$

with

$$\left\{Q^{a},\chi_{i}\right\}\approx\left\{\overline{Q}_{a},\chi_{i}\right\}\approx\left\{H,\chi_{i}\right\}\approx\left\{\chi_{i},\chi_{j}\right\}\approx0.$$
(2.20)

Details of this computation can be found in the Appendix. The supercharges Q^a and \overline{Q}_b in (2.17) and the Hamiltonian *H* in (2.18) describe the $\mathcal{N}=2M$ supersymmetric su(n) spin-Calogero model.

For $\mathcal{N}=4$ it essentially coincides with the osp(4|2) supersymmetric mechanics constructed in [16,17]. However, there are a few differences:

- The Hamiltonian (2.18) has no interaction for the center-ofmass coordinate $X = \sum_i x^i$. Correspondingly, the supercharges (2.17) do not include certain terms which appeared in [16,17].
- Working at the Hamiltonian level, we may keep the su(n) generators ℓ_{ij} unspecified. Precisely this enables the minimal realization (2.5) with a minimal number of auxiliary variables v_i , \bar{v}_i . At the Lagrangian level this corresponds to using (2, 4, 2) supermultiplets for the auxiliary bosonic superfields instead of (4, 4, 0) superfields as in [16,17].

Now we are ready to reduce our $\mathcal{N}=2M$ su(n) spin-Calogero model to a genuine $\mathcal{N}=2M$ Calogero model.

2.3. *N*-extended supersymmetric (no-spin) Calogero models

As we can see from the previous subsection, the supersymmetric analogs (2.19) of the purely bosonic constraints (2.3) appear automatically. These constraints generate n-1 local U(1) transformations³ of the variables { v_i , \bar{v}_i , ρ_{ij}^a , $\bar{\rho}_{ija}$ }. In terms of the 2*n* polar variables r_i and ϕ_i defined in (2.6), the constraints (2.19) can be easily resolved as

$$r_k^2 \approx r_n^2 + \Pi_{kk} - \Pi_{nn}$$
 for $k = 1, ..., n-1$. (2.21)

After fixing the residual gauge freedom as

$$\phi_1 \approx \phi_2 \approx \ldots \approx \phi_{n-1} \approx 0$$
, (2.22)

we obtain the supercharges and Hamiltonian which still obey the $\mathcal{N}=2M$ superalgebra (2.11) and contain only the surviving pair (r_n, ϕ_n) of the originally 2n "angular" variables. One may check that the supercharges Q^a and \overline{Q}_b and the Hamiltonian H, with the generators ℓ_{ij} replaced by $\hat{\ell}_{ij}$ and with the constraints (2.21) and (2.22) taken into account, perfectly commute with $r_n^2 - \Pi_{nn}$. Thus, the final step of the reduction is to impose the constraint

$$r_n^2 - \Pi_{nn} \approx \text{const} =: g$$
 (2.23)

and to fix the remaining U(1) gauge symmetry via

$$\phi_n \approx 0. \tag{2.24}$$

The previous two relations are the supersymmetric analogs of (2.10). We conclude that the full set of the reduction constraints reads

$$r_i^2 \approx g + \Pi_{ii}$$
 and $\phi_i \approx 0$ for $i = 1, \dots, n$. (2.25)

With these constraints taken into account, our supercharges Q^a and \overline{Q}_b and the Hamiltonian *H* acquire the form

$$\begin{split} \widehat{Q}^{a} &= \sum_{i=1}^{n} p_{i} \psi_{i}^{a} - i \sum_{i \neq j}^{n} \frac{\left(\sqrt{g + \Pi_{ii}} \sqrt{g + \Pi_{jj}} - \Pi_{ij}\right) \rho_{ji}^{a}}{x^{i} - x^{j}} ,\\ \widehat{\overline{Q}}_{b} &= \sum_{i=1}^{n} p_{i} \bar{\psi}_{ib} + i \sum_{i \neq j}^{n} \frac{\bar{\rho}_{ijb} \left(\sqrt{g + \Pi_{ii}} \sqrt{g + \Pi_{jj}} - \Pi_{ji}\right)}{x^{i} - x^{j}} ,\\ \widehat{H} &= \frac{1}{2} \sum_{i=1}^{n} p_{i}^{2} \\ &+ \frac{1}{2} \sum_{i \neq j}^{n} \frac{\left(\sqrt{g + \Pi_{ii}} \sqrt{g + \Pi_{jj}} - \Pi_{ij}\right) \left(\sqrt{g + \Pi_{ii}} \sqrt{g + \Pi_{jj}} - \Pi_{ji}\right)}{(x^{i} - x^{j})^{2}} . \end{split}$$
(2.26)

It is matter of quite lengthy and tedious calculations to check that these supercharges and Hamiltonian form an $\mathcal{N}=2M$ superalgebra (2.11). The strategy is given in the Appendix. The main complication arises from the expressions $\sqrt{g + \Pi_{ii}}$ present in the supercharges and the Hamiltonian. Due to the nilpotent nature of Π_{ij} , the series expansion eventually terminates, but even in the two-particle case with $\mathcal{N}=4$ supersymmetry we encounter a lengthy expression,

$$\sqrt{g + \Pi_{11}} = \sqrt{g} \left(1 + \frac{1}{2g} \Pi_{11} - \frac{1}{8g^2} \Pi_{11}^2 + \frac{1}{16g^3} \Pi_{11}^3 - \frac{5}{128g^4} \Pi_{11}^4 \right).$$
(2.27)

For *n* particles the series will end with a term proportional to $(\Pi_{ii})^{\mathcal{N}(n-1)}$. Clearly, these terms will generate higher-degree monomials in the fermions, both for the supercharges and for the

² The system with the Hamiltonian (2.18) and with $\ell_{ij} = 0$ has been previously considered in [20].

³ Due to the relation $\sum_{i=1}^{n} \chi_i = 0$ we have only n-1 independent constraints.

Hamiltonian. We can only speculate that the dread of such complexities impeded an earlier discovery of genuine $\mathcal{N}=4$ Calogero models.

2.4. Simplest example: $\mathcal{N}=2$ supersymmetric two-particle Calogero model

For $\mathcal{N}=2$ supersymmetry one has to put M = 1 in the expressions (2.26) for the supercharges and Hamiltonian. This somewhat reduces their complexity compared to the $\mathcal{N}=4$ case, but the real simplification occurs for two particles. Indeed, for n=2 we get

$$\Pi_{22} = -\Pi_{11} \quad \text{and} \quad \Pi_{11}^3 \equiv 0$$

$$\Rightarrow \quad \sqrt{g + \Pi_{11}} \sqrt{g - \Pi_{11}} = \left(g - \frac{1}{2g}\Pi_{11}^2\right) \quad \text{for} \quad g \neq 0.$$
(2.28)

Moreover, the term Π_{11}^2 is of the maximal possible power in the ρ and $\bar{\rho}$ fermions and, therefore, disappears from the supercharges. Thus, we are left with

$$\widehat{Q}_{(2)} = \sum_{i=1}^{2} p_{i}\psi_{i} - i\sum_{i\neq j}^{2} \frac{(g - \Pi_{ij})\rho_{ji}}{x^{i} - x^{j}} \quad \text{and}
\widehat{\overline{Q}}_{(2)} = \sum_{i=1}^{2} p_{i}\bar{\psi}_{i} + i\sum_{i\neq j}^{2} \frac{\bar{\rho}_{ij}(g - \Pi_{ji})}{x^{i} - x^{j}}, \qquad (2.29)$$

which have the standard structure – linear and cubic in the fermions. The Hamiltonian $\hat{H}_{(2)}$ reduces to

$$\widehat{H}_{(2)} = \frac{1}{2} \sum_{i=1}^{2} p_i^2 + \frac{g^2 - \Pi_{11}^2 - g(\Pi_{12} + \Pi_{21}) + \Pi_{12}\Pi_{21}}{\left(x^1 - x^2\right)^2} , \quad (2.30)$$

with the explicit expressions

$$\Pi_{11} = \rho_{12}\bar{\rho}_{21} + \bar{\rho}_{12}\rho_{21} ,$$

$$\Pi_{12} = (\psi_1 - \psi_2)\bar{\rho}_{12} + (\bar{\psi}_1 - \bar{\psi}_2)\rho_{12} ,$$

$$\Pi_{21} = (\psi_2 - \psi_1)\bar{\rho}_{21} + (\bar{\psi}_2 - \bar{\psi}_1)\rho_{21} .$$
(2.31)

This $\mathcal{N}=2$ supersymmetric two-particle Calogero model has been previously constructed and analyzed in [16] (for details see the review [9]). This demonstrates that our approach perfectly reproduces the unique known $\mathcal{N}=2$ example.

3. Conclusion

We propose a novel N-extended supersymmetric su(n) spin-Calogero model as a direct supersymmetrization of the bosonic su(n) model [12]. In the case of N=4 supersymmetry, our model resembles the one constructed in [16,17]. However, there are two main differences:

- the center of mass is free
- the *su*(*n*) generators are not specified in a particular realization.

Thanks to these features, we were able to generalize the reduction procedure to the no-spin Calogero model from $\mathcal{N}=4$ supersymmetry to any number $\mathcal{N}=2M$ of supersymmetries. This led to the discovery of a genuine $\mathcal{N}=2M$ supersymmetric rational Calogero model for any number of particles.

Our models belong to same class which was proposed in [16,17]. Its main features are

- a huge number of fermionic coordinates, namely Nn^2 in number rather than the Nn to be expected
- the supercharges and the Hamiltonian contain terms which a fermionic power much larger than three.

Clearly, these features merit a more careful and detailed analysis. The following further developments come to mind:

- a superspace description of the constructed models, at least for $\mathcal{N}=2$ and $\mathcal{N}=4$ supersymmetry, presumably with nonlinear chiral supermultiplets
- an extension to the Calogero–Sutherland inverse-sine-square model
- an extension to the Euler–Calogero–Moser system [11] and its reduction to the goldfish system [13], yielding a supersymmetric goldfish model upon reduction, to be compared with recent results from [21].

Acknowledgements

We are grateful to Sergey Fedoruk for stimulating discussions. This work was partially supported by the Heisenberg–Landau program. The work of S.K. was partially supported by Russian Science Foundation grant 14-11-00598, the one of A.S. by RFBR grants 18-02-01046 and 18-52-05002 Arm-a. This article is based upon work from COST Action MP1405 QSPACE, supported by COST (European Cooperation in Science and Technology).

Appendix A. Details of the calculations

We check that the \mathcal{N} supercharges Q^a and \overline{Q}_b in (2.17) generate the $\mathcal{N}=2M$ superalgebra (2.11) with the Hamiltonian (2.18) modulo the first-class constraints (2.19). The best way to perform this calculation is to introduce the composite object

$$L_{ij} = \ell_{ij} + \Pi_{ij} \,. \tag{A.1}$$

Like ℓ_{ij} and Π_{ij} , their sums also form an su(n) algebra,

$$L_{ij}, L_{km} \} = i \left(\delta_{im} L_{kj} - \delta_{kj} L_{im} \right) , \qquad (A.2)$$

and they Poisson-commute with the fermions exactly as Π_{ij} in (2.16),

$$\begin{split} \left\{ L_{ij}, \psi_k^a \right\} &= i \left(\delta_{ik} - \delta_{jk} \right) \rho_{ij}^a , \\ \left\{ L_{ij}, \rho_{km}^a \right\} &= -i \, \delta_{im} \delta_{jk} \left(\psi_i^a - \psi_j^a \right) - i \delta_{jk} \rho_{im}^a + i \delta_{im} \rho_{kj}^a , \\ \left\{ L_{ij}, \bar{\psi}_{ka} \right\} &= i \left(\delta_{ik} - \delta_{jk} \right) \bar{\rho}_{ija} , \\ \left\{ L_{ij}, \bar{\rho}_{kma} \right\} &= -i \, \delta_{im} \delta_{jk} \left(\bar{\psi}_{ia} - \bar{\psi}_{ja} \right) - i \delta_{jk} \bar{\rho}_{ima} + i \delta_{im} \bar{\rho}_{kja} . \end{split}$$

$$(A 3)$$

It should be clear now that the closing of the superalgebra (2.11) for the supercharges

$$Q^{a} = \sum_{i=1}^{n} p_{i} \psi_{i}^{a} + i \sum_{i \neq j}^{n} \frac{L_{ij} \rho_{ji}^{a}}{x^{i} - x^{j}} \quad \text{and}$$

$$\overline{Q}_{b} = \sum_{i=1}^{n} p_{i} \overline{\psi}_{ib} - i \sum_{i \neq j}^{n} \frac{\overline{\rho}_{ijb} L_{ji}}{x^{i} - x^{j}} \quad (A.4)$$

does not depend on the number of particles or the number of supersymmetries. It is based on the Poisson brackets (A.2) and (A.3) and the basic brackets (2.2) and (2.13), i.e.

$$\{x^{i}, p_{j}\} = \delta^{i}_{j} \quad \text{and} \quad \{\psi^{a}_{i}, \bar{\psi}_{jb}\} = -i \,\delta^{a}_{b} \delta_{ij} ,$$

$$\{\rho^{a}_{ij}, \bar{\rho}_{kmb}\} = -i \,\delta^{a}_{b} \delta_{im} \delta_{jk} .$$
(A.5)

The computation makes repeated use of the all-important identity

$$\frac{1}{(x^{i} - x^{j})(x^{i} - x^{k})} + \frac{1}{(x^{j} - x^{i})(x^{j} - x^{k})} + \frac{1}{(x^{k} - x^{i})(x^{k} - x^{j})} = 0$$

for $i \neq j \neq k$. (A.6)

Direct calculation then yields

$$\{Q^{a}, Q^{b}\} = i \sum_{i \neq j} \frac{\rho_{ij}^{a} \rho_{ji}^{b}}{(x^{i} - x^{j})^{2}} (L_{ii} - L_{jj}),$$

$$\{\overline{Q}_{a}, \overline{Q}_{b}\} = i \sum_{i \neq j} \frac{\overline{\rho}_{ija} \overline{\rho}_{jib}}{(x^{i} - x^{j})^{2}} (L_{ii} - L_{jj}), \qquad (A.7)$$

$$\left\{Q^{a}, \overline{Q}_{b}\right\} = -2i\delta_{b}^{a}H + i\sum_{i\neq j}\frac{\rho_{ij}^{a}\bar{\rho}_{jib}}{(x^{i}-x^{j})^{2}}\left(L_{ii}-L_{jj}\right),$$

where

$$H = \frac{1}{2} \sum_{i=1}^{n} p_i^2 + \frac{1}{2} \sum_{i \neq j}^{n} \frac{L_{ij} L_{ji}}{\left(x^i - x^j\right)^2} \,. \tag{A.8}$$

Clearly, imposing the constraints (2.19),

$$\chi_i = \ell_{ii} + \Pi_{ii} = L_{ii} \approx 0 \quad \forall i \quad (\text{no sum}), \qquad (A.9)$$

closes the superalgebra (A.7). Since $\sum_{i} L_{ii} = 0$ reduces u(n) to su(n), the constraints (A.9) cannot be relaxed to $L_{ii} \approx \alpha \neq 0$.

With the same strategy we can check that the supercharges \widehat{Q}^a and $\overline{\widehat{Q}}_a$ and the Hamiltonian \widehat{H} in (2.26) form the same superalgebra (2.11). We do not need to go inside the objects Π_{ij} or $\sqrt{g + \Pi_{ii}}$. Instead, we directly employ the Poisson brackets of the composites,

$$\begin{aligned} \{\Pi_{ij}, \Pi_{km}\} &= i(\delta_{im}\Pi_{kj} - \delta_{kj}\Pi_{im}), \\ \{\Pi_{ij}, \sqrt{g + \Pi_{kk}}\} &= \frac{i}{2\sqrt{g + \Pi_{kk}}}(\delta_{ik}\Pi_{kj} - \delta_{kj}\Pi_{ik}), \\ \{L_{ij}, \psi^a_k\} &= i(\delta_{ik} - \delta_{jk})\rho^a_{ij}, \\ \{L_{ij}, \rho^a_{km}\} &= -i\delta_{im}\delta_{jk}(\psi^a_i - \psi^a_j) - i\delta_{jk}\rho^a_{im} + i\delta_{im}\rho^a_{kj}, \\ \{L_{ij}, \bar{\psi}_{ka}\} &= i(\delta_{ik} - \delta_{jk})\bar{\rho}_{ija}, \\ \{L_{ij}, \bar{\rho}_{kma}\} &= -i\delta_{im}\delta_{jk}(\bar{\psi}_{ia} - \bar{\psi}_{ja}) - i\delta_{jk}\bar{\rho}_{ima} + i\delta_{im}\bar{\rho}_{kja}, \end{aligned}$$

$$\begin{aligned} \left\{ \sqrt{g} + \Pi_{ii}, \psi_k^a \right\} &= 0 ,\\ \left\{ \sqrt{g} + \Pi_{ii}, \rho_{km}^a \right\} &= -\frac{\mathrm{i}}{2\sqrt{g} + \Pi_{ii}} \left(\delta_{ik} \rho_{im}^a - \delta_{im} \rho_{ki}^a \right) ,\\ \left\{ \sqrt{g} + \Pi_{ii}, \bar{\psi}_{ka} \right\} &= 0 ,\\ \left\{ \sqrt{g} + \Pi_{ii}, \bar{\rho}_{kma} \right\} &= -\frac{\mathrm{i}}{2\sqrt{g} + \Pi_{ii}} \left(\delta_{ik} \bar{\rho}_{ima} - \delta_{im} \bar{\rho}_{ija} \right) . \end{aligned}$$

$$(A.10)$$

This drastically simplifies the evaluation and, by using again (A.6), one may readily convince oneself that indeed

$$\{\widehat{Q}^{a}, \widehat{\overline{Q}}_{b}\} = -2i\delta_{b}^{a}\widehat{H}$$
 and $\{\widehat{Q}^{a}, \widehat{Q}^{b}\} = \{\widehat{\overline{Q}}_{a}, \widehat{\overline{Q}}_{b}\} = 0.$
(A.11)

References

- [1] F. Calogero, J. Math. Phys. 10 (1969) 2191 10 (1969) 2179 12 (1971) 419.
- [2] D.Z. Freedman, P.F. Mende, Nucl. Phys. B 344 (1990) 317.
- [3] N. Wyllard, J. Math. Phys. 41 (2000) 2826, arXiv:hep-th/9910160.
- [4] S. Bellucci, A. Galajinsky, E. Latini, Phys. Rev. D 71 (2005) 044023, arXiv:hepth/0411232.
- [5] A. Galajinsky, O. Lechtenfeld, K. Polovnikov, J. High Energy Phys. 0711 (2007) 008, arXiv:0708.1075 [hep-th].
- [6] S. Bellucci, S. Krivonos, A. Sutulin, Nucl. Phys. B 805 (2008) 24, arXiv:0805.3480 [hep-th].
- [7] S. Fedoruk, E. Ivanov, O. Lechtenfeld, Phys. Rev. D 79 (2009) 105015, arXiv: 0812.4276 [hep-th].
- [8] S. Krivonos, O. Lechtenfeld, J. High Energy Phys. 1102 (2011) 042, arXiv:1012. 4639 [hep-th].
- [9] S. Fedoruk, E. Ivanov, O. Lechtenfeld, J. Phys. A 45 (2012) 173001, arXiv:1112. 1947 [hep-th].
- [10] D. Kazhdan, B. Konstant, A. Sternberg, Commun. Pure Appl. Math. 31 (1978) 481.
- [11] S. Wojciechowski, Phys. Lett. A 111 (1985) 101.
- [12] J. Gibbons, T. Hermsen, Physica D 11 (1984) 337.
- [13] J. Arnlind, M. Bordemann, J. Hoppe, C. Lee, Lett. Math. Phys. 84 (2008) 89, arXiv:math-ph/0702091.
- [14] A. Polychronakos, Phys. Lett. B 266 (1991) 29.
- [15] A. Polychronakos, J. Phys. A 39 (2006) 12793, arXiv:hep-th/0607033.
- [16] S. Fedoruk, E. Ivanov, O. Lechtenfeld, Phys. Rev. D 79 (2009) 105015, arXiv: 0812.4276 [hep-th].
- [17] S. Fedoruk, E. Ivanov, O. Lechtenfeld, S. Sidorov, J. High Energy Phys. 1804 (2018) 043, arXiv:1801.00206 [hep-th].
- [18] S. Fedoruk, E. Ivanov, O. Lechtenfeld, Phys. At. Nucl. 74 (2011) 870, arXiv:1001. 2536 [hep-th].
- [19] S. Fedoruk, E. Ivanov, J. High Energy Phys. 1611 (2016) 103, arXiv:1610.04202 [hep-th].
- [20] N.B. Copland, S.M. Ko, J.-H. Park, J. High Energy Phys. 1207 (2012) 076, arXiv: 1205.3869 [hep-th].
- [21] A. Galajinsky, J. High Energy Phys. 1804 (2018) 079, arXiv:1802.08011 [hep-th].